

rosetta



Philae Landing Preparations

A recall of what happened with a personal touch

ESTEC

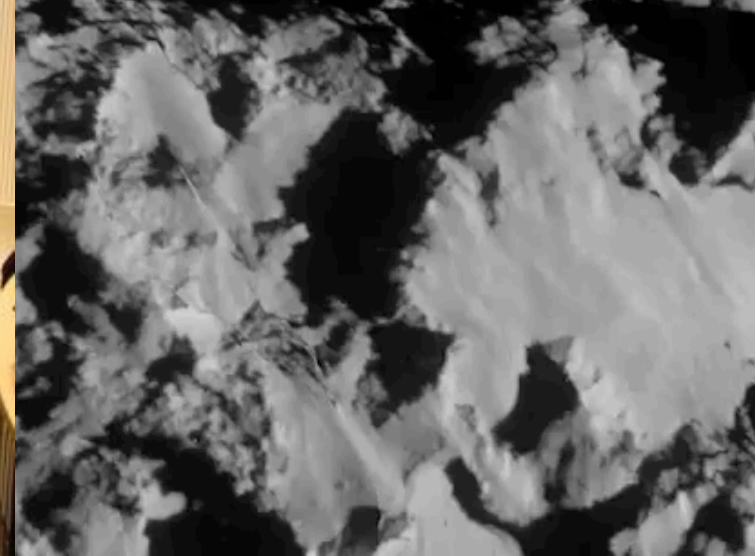
10 January 2018

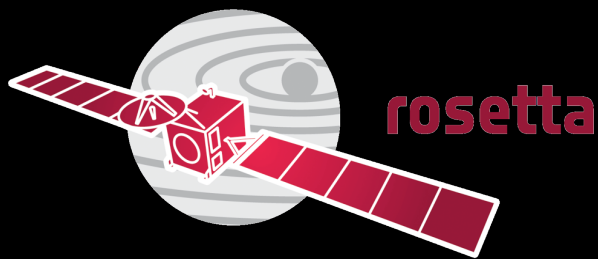
Joe.Zender@esa.int

Matt.Taylor@esa.int



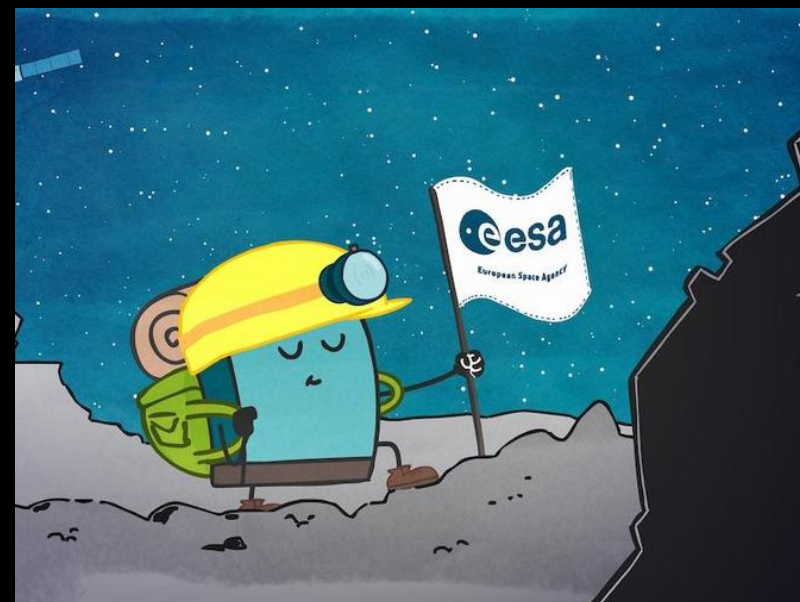
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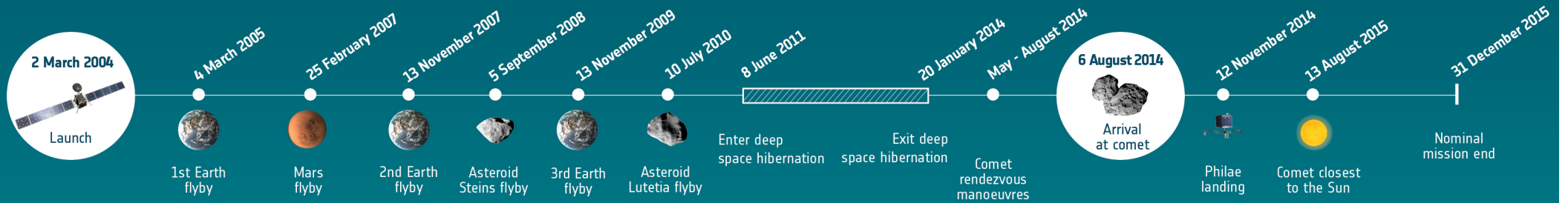
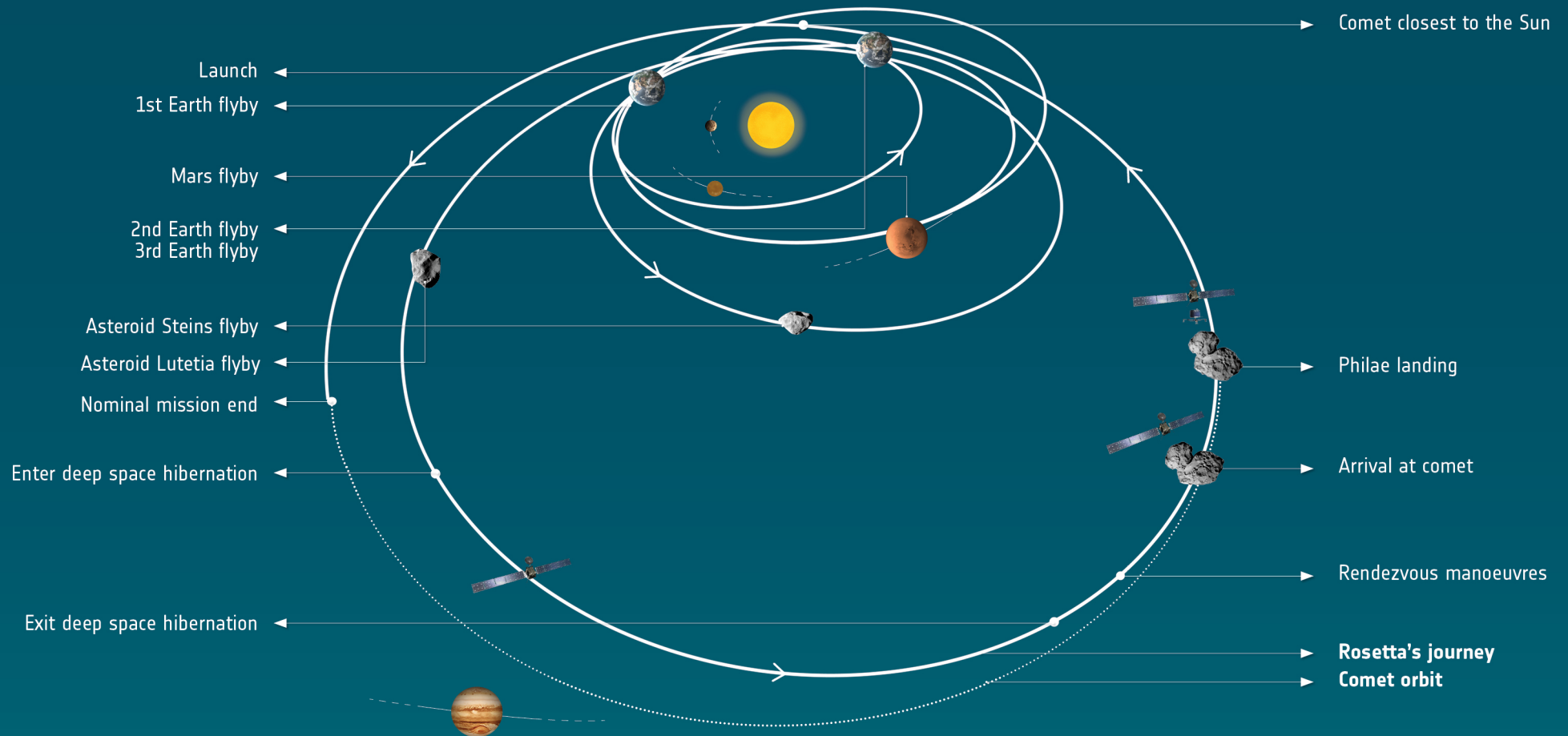


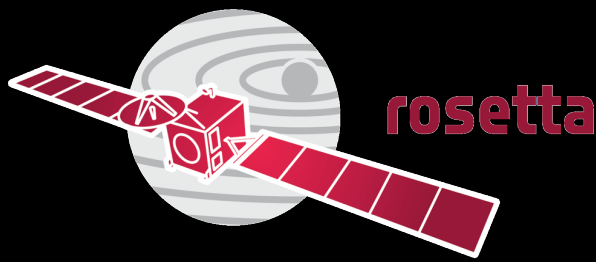
Outline

- The year 2014: remember what we did not know
- Our Schedule: step by step



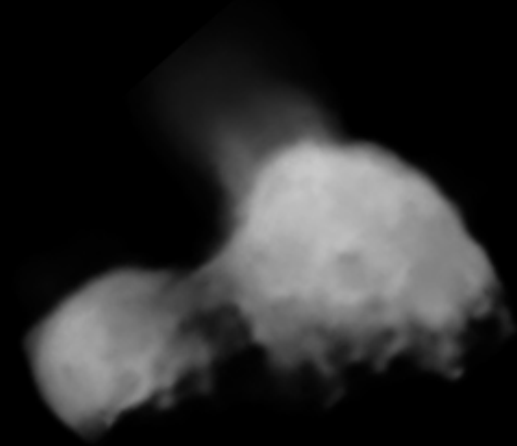
→ ROSETTA'S JOURNEY





Spacecraft Visits to Comets (imaged)

- The Halley Armada
Giotto, Vega 1 and 2, Suisei, Sakigake
- Deep Space 1 (Borrelly)
- Stardust (Wild 2)
- Deep Impact (Tempel 1)
- EPOXI (Hartley 2)



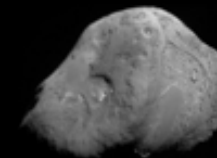
1P/Halley - $16 \times 8 \times 8$ km



81P/Wild 2
 $5.5 \times 4.0 \times 3.3$ km
Stardust, 2004



19P/Borrelly
 8×4 km
Deep Space 1, 2001

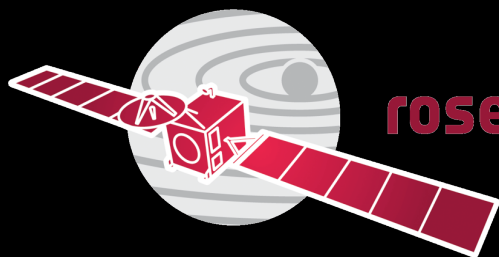


9P/Tempel 1
 7.6×4.9 km
Deep Impact, 2005



103P/Hartley 2
 2.2×0.5 km
Deep Impact, 2010

Fly by's - 100's km
10's km/s



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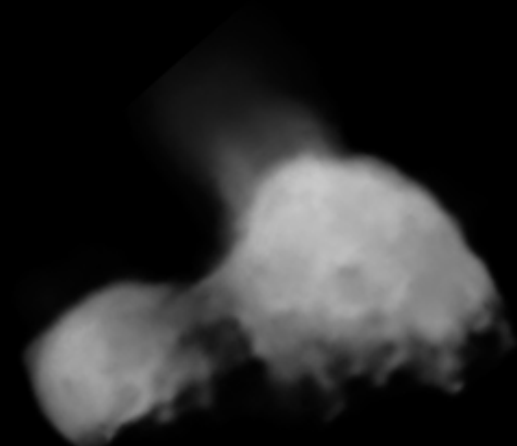
1P/Halley: Highly active, low albedo, relatively little geological information about the surface

19P/Borrelly: Diverse geology, different types of terrain, no ice found on surface!

81P/Wild: Rugged terrain, impact craters ?

9P/Tempel 1: Diverse terrain, primordial layers found?, impact craters ?, very little ice found on surface

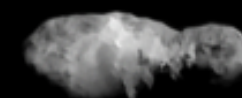
103P/Hartley 2: Hyperactive, diverse terrain, extreme shape, ice blocks (cm-dm sized) emitted from nucleus



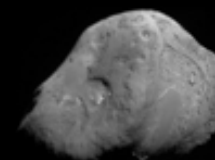
1P/Halley - $16 \times 8 \times 8$ km



81P/Wild 2
 $5.5 \times 4.0 \times 3.3$ km
Stardust, 2004



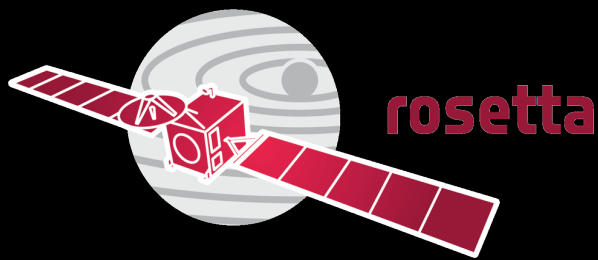
19P/Borrelly
 8×4 km
Deep Space 1, 2001



9P/Tempel 1
 7.6×4.9 km
Deep Impact, 2005



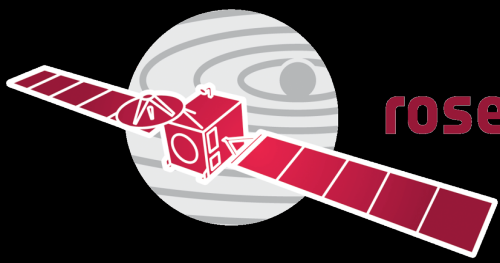
103P/Hartley 2
 2.2×0.5 km
Deep Impact, 2010



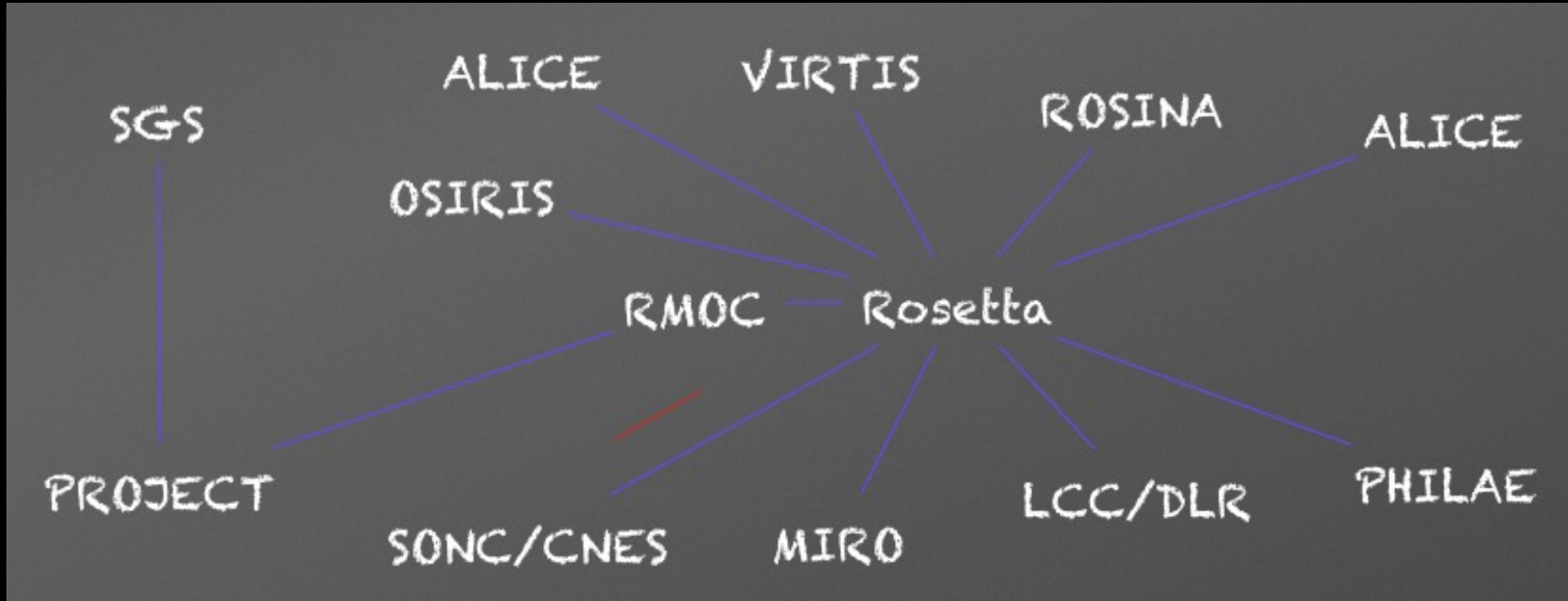
My task:

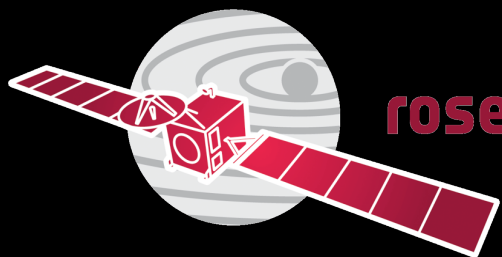
Interface between the orbiter teams and the landing teams in all aspects of the landing site selection preparation.

“Identify crucial tasks and get them done.”



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Step-by-step



Collect information and talk to the teams:

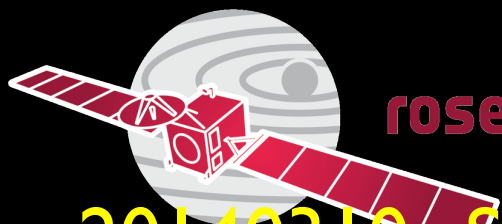
20140206: telecon with OSIRIS camera team

20140207: telecon with ROSINA ion/neutral spectrometer team

20140207: telecon with VIRTIS UV-spectrometer team

20140214: telecon with MIRO IR-spectrometer team

20140218: telecon with ALICE UV-spectrometer team



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Step-by-step

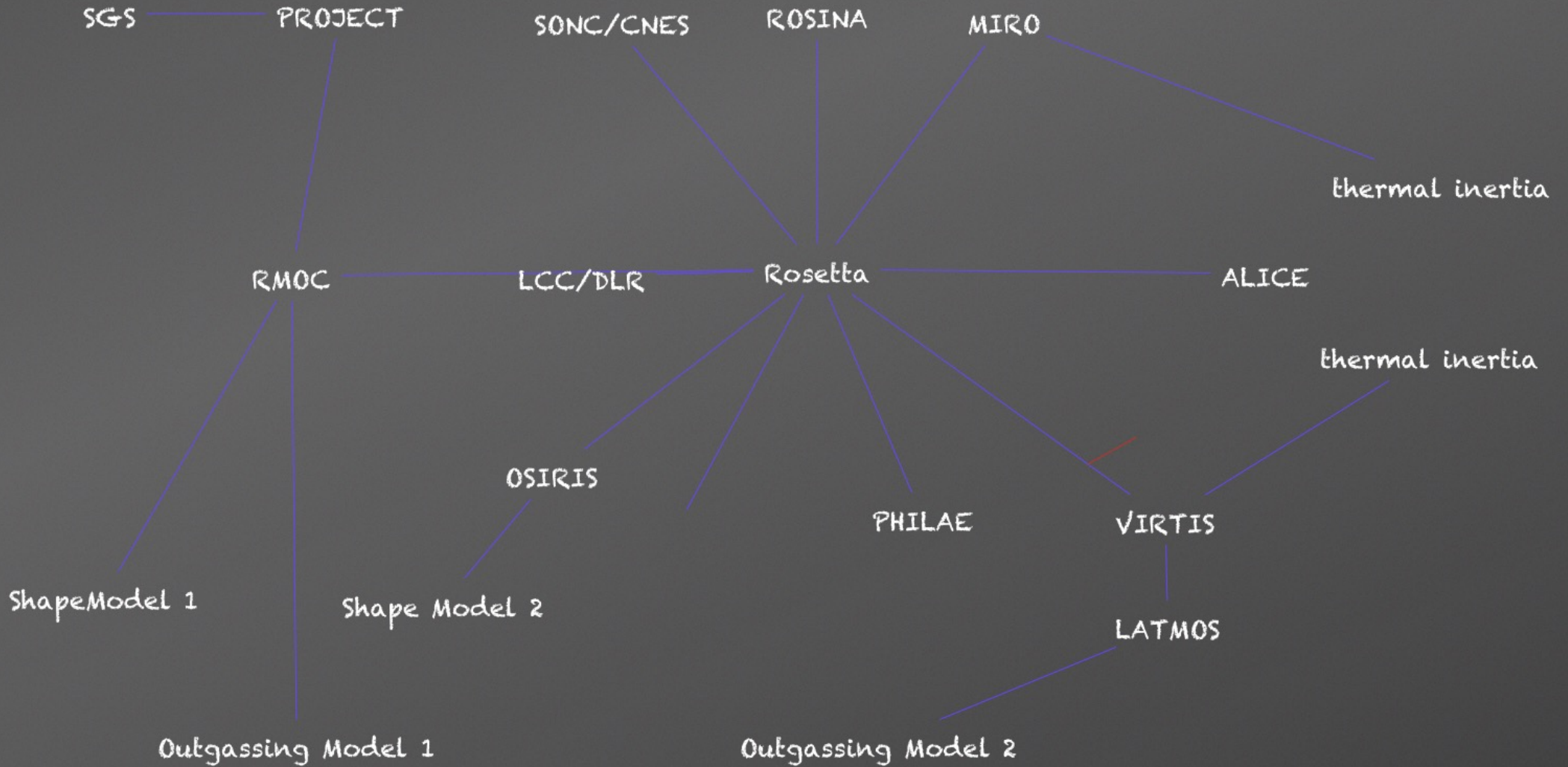
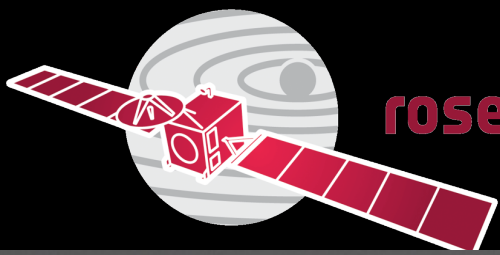


20140310: Science Working Team meeting

- 2 outgassing models (RMOC, LATMOS)
- who will produce the thermal model? 2 candidates: MIRO or VIRTIS?
 - MIRO indicates that they miss the comet dark side shape information: it is not clear what the impact not the thermal model is.
 - MIRO asks, if the SONC needs the surface gas velocity or the surface gas velocity profile? Action item for SONC.



Date/Time	Room	Topic	Participants
Monday, 14:30	H168	RLLS Validation Test discussion	SONC RLGS SGS
Monday, 16:00	H168	RLLS Validation Test: test input data and	OSIRIS RLGS MOC SONC
TBC		OSIRIS derived SPICE kernels: needs and definition	BG,MAH
Monday 18:00	H168	Comet Surface Temperature Model (CSTM): input, model, responsibility	MIRO VIRTIS SONC RLGS
Wednesday 16:00 (TBC)		LATMOS interfaces	LATMOS SONC RLGS VIRTIS
Wednesday, 14:00		Coordinate Reference Frames	SGS MOC OSIRIS

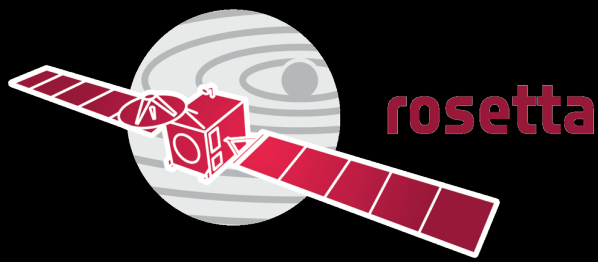


Description	Action	Due Date
<p><u>Model properties</u></p> <p>It was clarified that the RMOC model always solves for full (complex) rotation parameters. If the rotation is simple, this may be reflected in the parameters, but those parameters are not explicitly calculated. The origin of the coordinate system of the RMOC model is the centre of mass of the comet. The coordinate system may change slightly when the quality of the model is increased with increasing image resolution, but it will not be change arbitrarily when the model is updated. Flight dynamics pointed out that landmark positions are provided regularly both as Cartesian body-fixed coordinates and as pixel positions in images.</p> <p>The OSIRIS model will try to solve for simple rotation first and go for complex rotation if the residuals for simple rotation are larger than the error bars. In the OSIRIS shape model, the centre of the reference frame is the geometric centre of the shape. The difference between the centres in the two reference frames is expected to be insignificant (to be tested by OSIRIS).</p> <p>The OSIRIS model will use the rotation parameters from the RMOC model as a starting point for reconstruction. Those parameters may be (slightly) modified in the optimization process.</p>		
<p><u>Frame transformation between RMOC model and OSIRIS model:</u></p>		

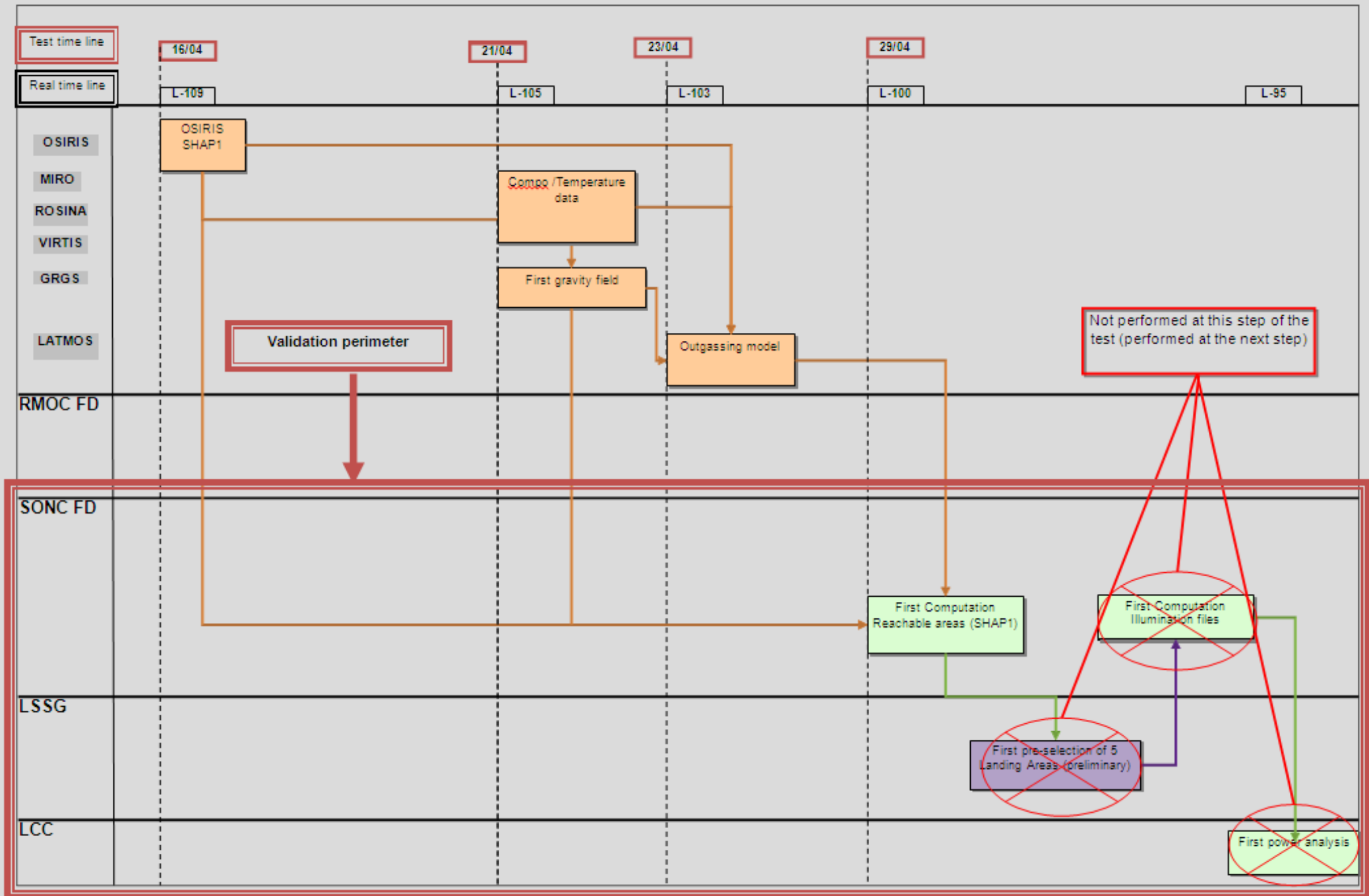


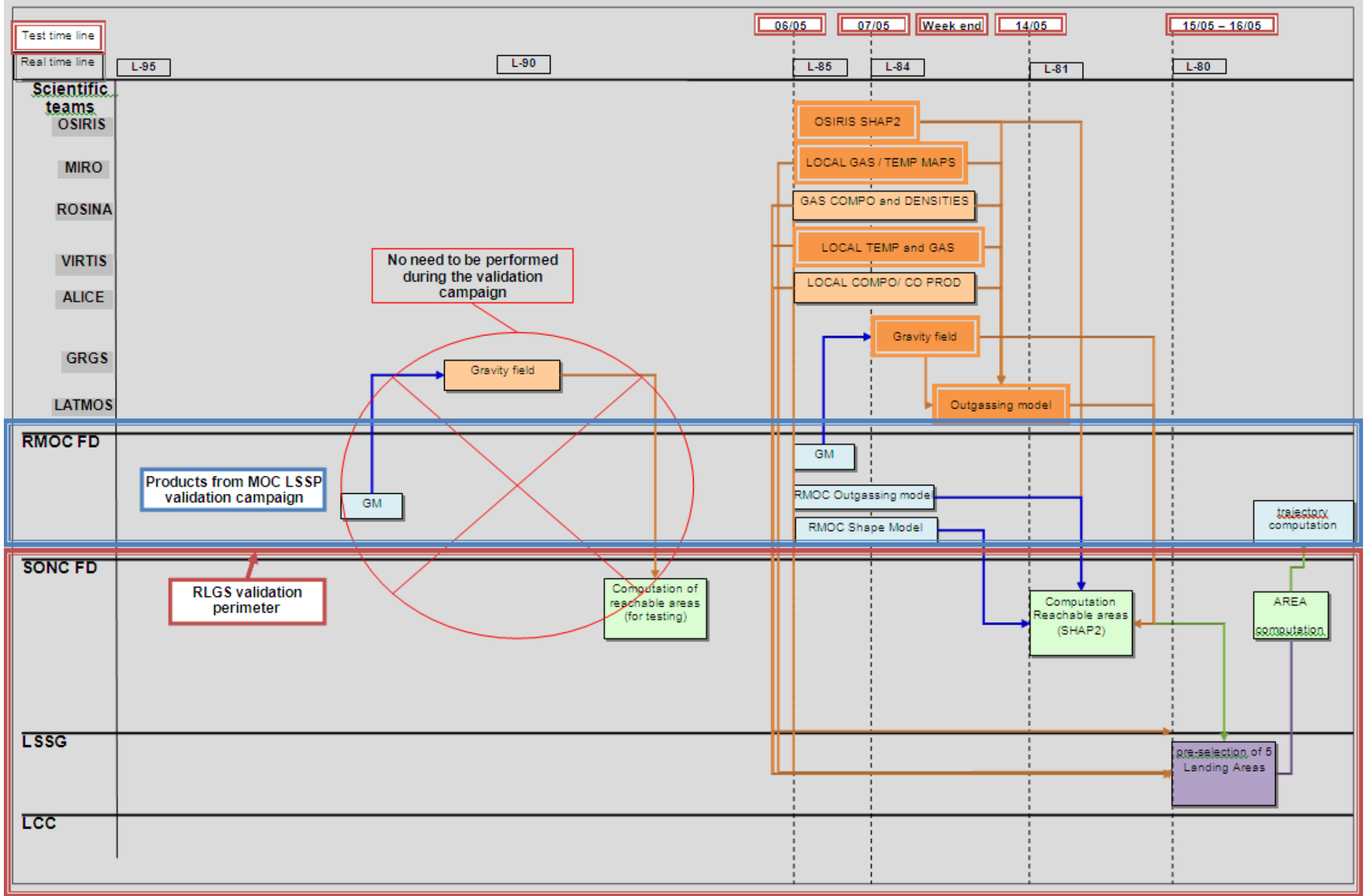
Navigation Camera NAVCAM1 20140608

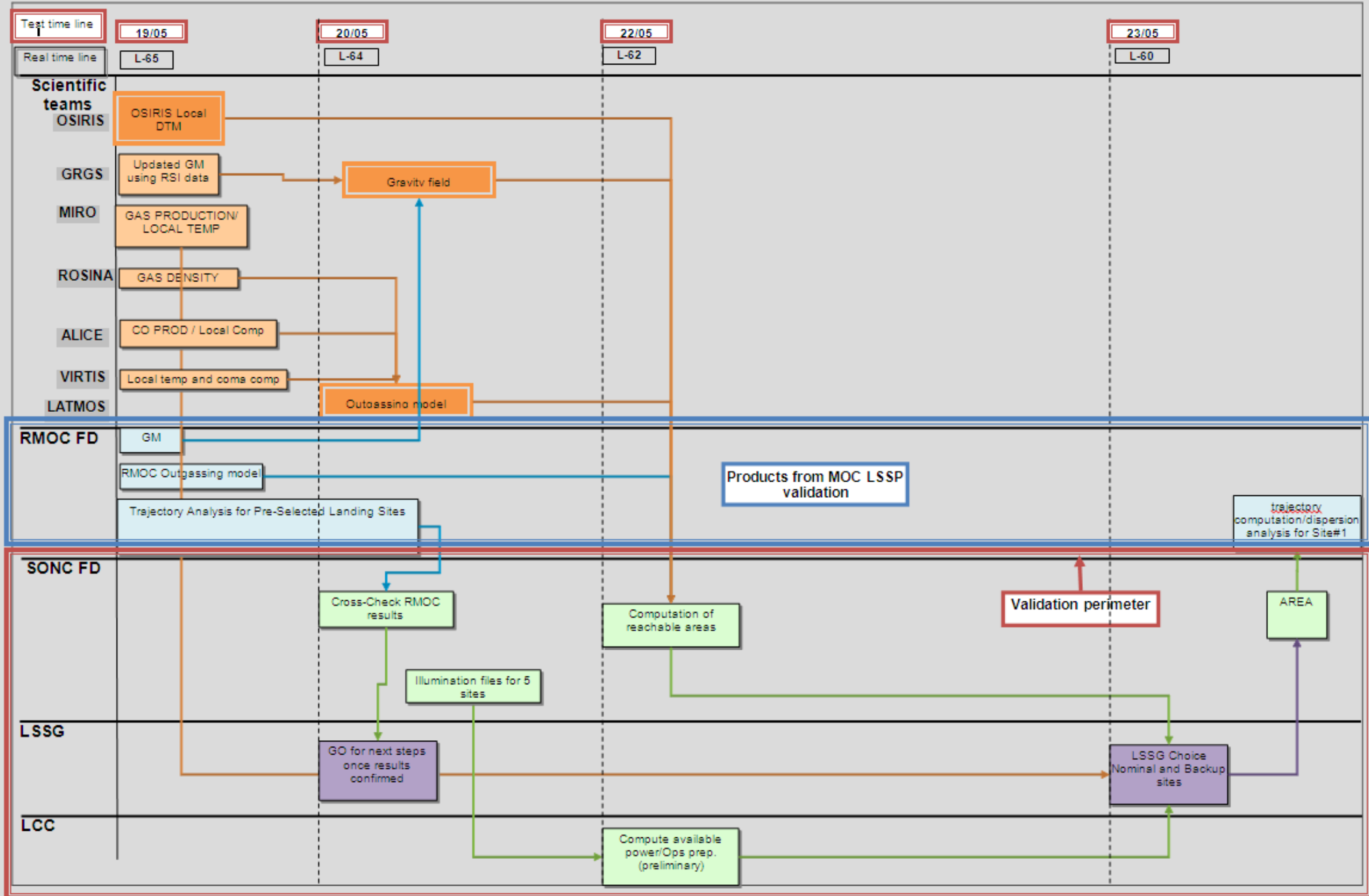


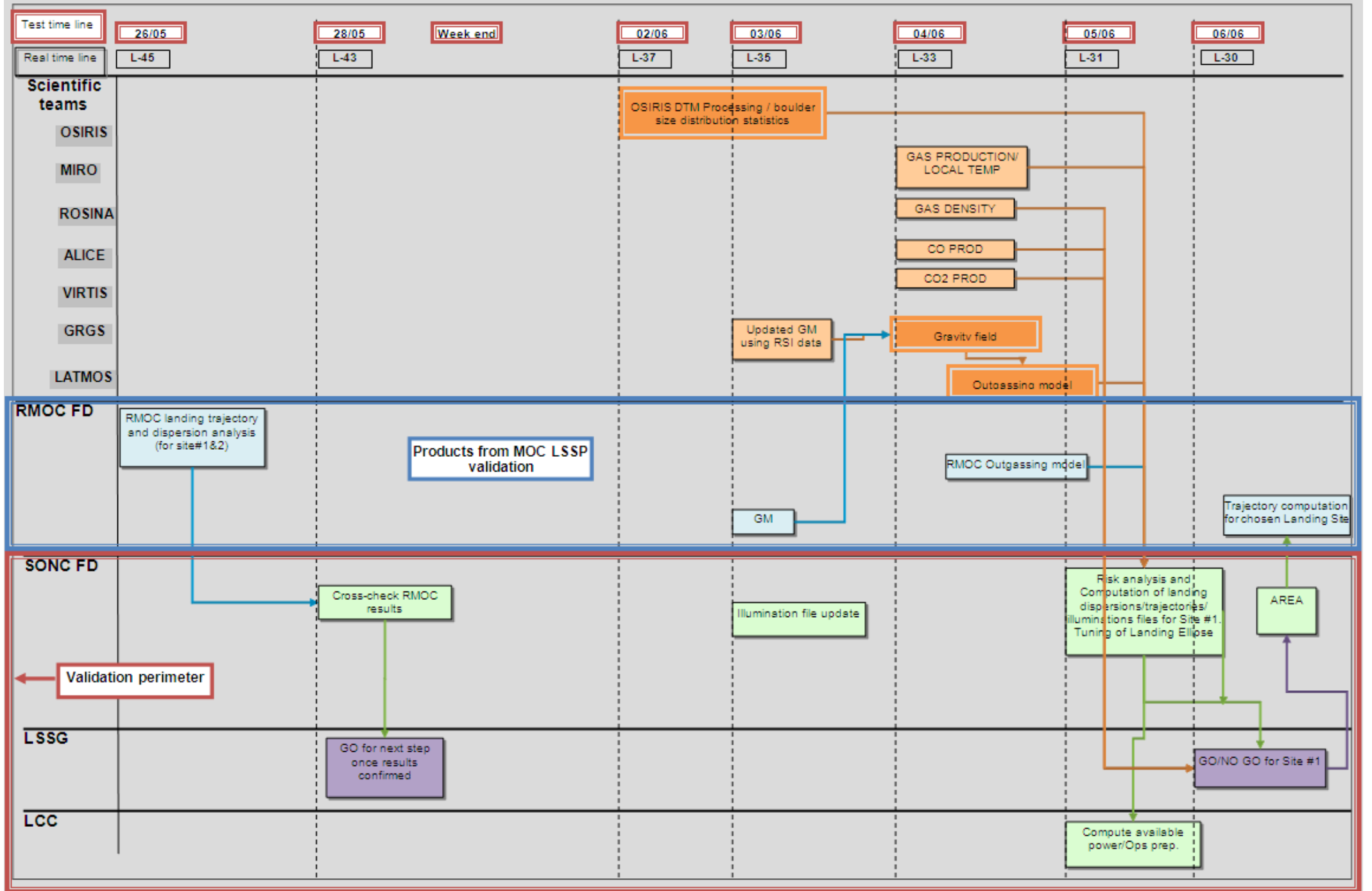


The Plan the Test Plan first ... as seen by SONC in
March 2014









Road Map



- Clarification about the usable inputs (RMOC & PI products) => from now to end of March
 - ▶ Usable products have to be prepared as soon as possible in order to ensure their availability for other providers who need them, and to ensure their delivery when expected for the test,
- Finalization of LSSP test with RMOC => end of March
- Test plan update/finalization => end of March
- “Warm up” test (if possible) => during April
- Test Readiness Review => end April
- LSSP phase 1 validation => 06/05 to 16/05
- LSSP phase 2 validation => 19/05 to 23/05
- LSSP phase 3 validation => 26/05 to 06/06
- Post Test Review => mid June
- ESA LSSP Readiness Meeting => 01/07



FMI



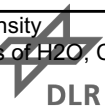
STFC



Products (wrt ICDs)	Classification	Producer	Product available for test?
Shape model "RS_GLOBAL_DTM_500m", "RS_GLOBAL_DTM_20m"	SHAP1 important; SHAP2 & SHAP2MIRO Mandatory	OSIRIS	
rotational parameters/model "RS_ROT_PARAM_500m", "RS_ROT_PARAM_20m", "RS_ROT_PARAM_LOCAL_DTM", "RS_ROT_PARAM_IMAGES_LS"	Mandatory	OSIRIS	
Local DTM (3m resolution) "RS_LOCAL_DTM"	Mandatory	OSIRIS	
Images high resolution "RS_IMAGES_LAND_SITE1", "RS_IMAGES_LAND_SITE2", ,	Mandatory	OSIRIS	
obstacle size distribution "RS_BOULDERS_LAND_SIT E1", "RS_BOULDERS_LAND_SIT E2"	Mandatory	OSIRIS	
Colormaps, composition "RS_GLOBAL_COMPOSITION", "RS_LOCAL_COMPOSITION"	Important	OSIRIS	
Albedo map "RS_ALBEDO_500m", "RS_ALBEDO_20m"	Valuable	OSIRIS	
Dust and gas coma monitoring "RS_MONITORING_DUST_date", "RS_MONITORING_GAS_date"	Valuable	OSIRIS	
Inner coma monitoring "RS_MONITORING_INNER_COMA_date »"	Valuable	OSIRIS	
"Far global density measurements of H2O, CO and CO2", "	Important	ROSINA	

Products (wrt ICDs)	Classification	Producer	Product available for test?
"Global density measurements of H2O, CO and CO2"	Important	ROSINA	
"Close local measurements of H2O, CO and CO2"	Important	ROSINA	
"Global measurements of total gas density"	Important	ROSINA	
"Close local measurements of total gas density"	Important	ROSINA	
"Total density and gas dynamics measurements"	Important	ROSINA	
"Global Surface compos"	Important	ALICE	
"Surface of Candidate Landing Sites"	Important	ALICE	
"Global CO Gas Production"	Important	ALICE	
"CO Gas Production at Candidate Landing Sites "	Important	ALICE	
"Surface composition"	Important	VIRTIS	
"Surface temperature predictions and thermal properties" (TI)	Mandatory	VIRTIS & MIRO	
"Coma composition"	Important	VIRTIS	
"Global H2O and CO gas production rates"	Mandatory	MIRO	
"Local gas production and variability"	Important	MIRO	
« Temperature predictions of nucleus »" (TI)	Mandatory	MIRO & VIRTIS	
"Temperature map of candidate landing sites"	Important	MIRO	
"Gas production, velocity, and variability of candidate landing sites"	Important	MIRO	
"Antenna temperatures during selected night-time limb crossings"	Important	MIRO → Osiris	

from Jens Biele, LCC, March 2014



Landing Site Selection: Status of essential Orbiter Instrument Products



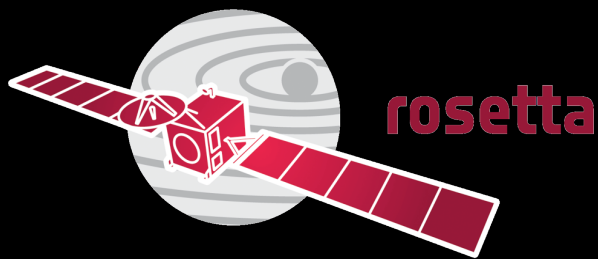
Product	Originator	Data Pipeline	Interface validation	E2E test preparation
20m Products	OSIRIS	in validation	Executed, awaiting confirmation, delta tests necessary	In preparation
3m Products	OSIRIS	In validation	Executed (ac)	In preparation
High-resolution images	OSIRIS	validated	Executed (ac)	Ready for test
List of boulders	OSIRIS	validated	Executed (ac)	Ready for test
Comet Surface Temp Model	VIRTIS/MIRO	In development	In preparation	In preparation
Gravity model	GRGS/CNES	No information	No information	No information
Outgassing model	LATMOS/CNES	No information	No information	No information

my own presentation from April 2014 at Steering Committee

Landing Site Selection: Status of important Orbiter Instrument Products (not part of E2E test in May 2014)



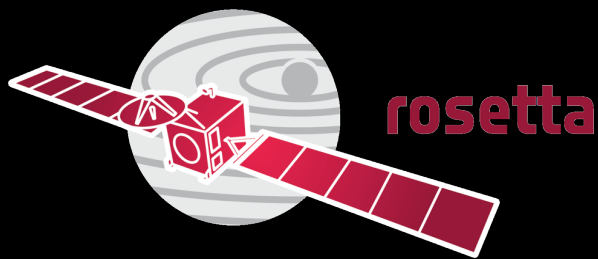
Product	Originator	Data Pipeline	Interface validation
H2O, CO and CO2 measurements	ROSINA	Validated	executed successfully
Gas density	ROSINA	Validated	Executed successfully
Surface composition (global/local)	ALICE	Validated	Ready for testing
CO gas production (global/local)	ALICE	Validated	Ready for testing
COMA composition	VIRTIS	validated	Executed successfully
Surface composition	VIRTIS	validated	Executed successfully
H2O, CO gas production	MIRO	In development	Executed successfully
Temperature maps (global/local)	MIRO	In development	Executed successfully
	MIRO	In development	Executed



To get all the model output, one needs
first: to plan and acquire the data
second: to downlink the data
and third: to have good data and success with the
computations.

and here is THE Plan ... as seen from Project ...

[open 01-RO-SGS-LI-1014_I_3_LSSP_InputProductsOverview_20140730](#)



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Juni 2014

gebaseerd op

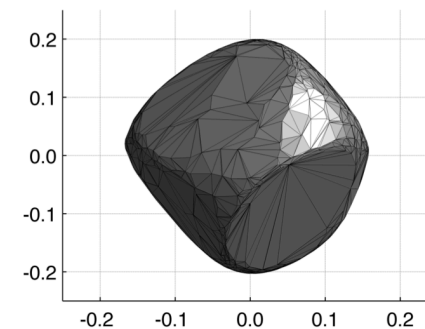
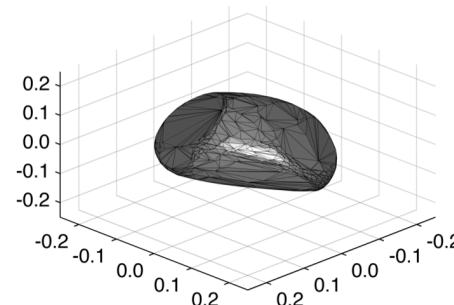
- > Optische technieken (lightcurves) uit telescoop waarnemingen
- > Thermische analyse en optische fotometrie
- > Spitzer thermische IR data

We verwachten (bijv. Op basis van Lowry et al, 2012)

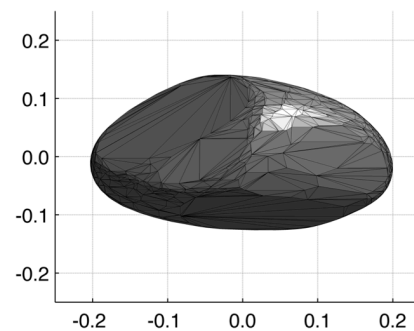
- > Axiale ratio's: $b / a = 1.239$, $c / a = 0,819$ (7% marge)
- > Spin rate: 12.76 uur
- > Pol: $\lambda = 78 (+ -10)^\circ$, $\beta = 58 (+ -10)$ graden
- > Nucleus fase verduistering: $G = 0,11 (+ -0,12)$, $H_r = 15,31 (+ - 0.07)$
- > Thermische inertie: $< 15 \text{ J} / \text{m}^2 / \text{K} / \text{sqrt}(\text{s})$

S. Lowry et al.: The Nucleus of Comet 67P/Churyumov-Gerasimenko

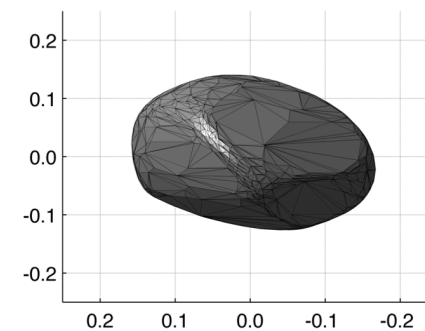
Pole Solution: $\lambda = 78$, $\beta = 58$



View along Z-axis

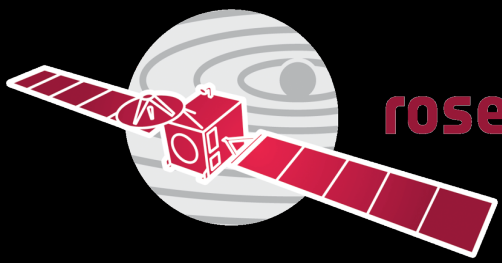


View along X-axis



View along Y-axis

Fig. 5. Best-fit convex triangular-facet shape model produced via lightcurve inversion using the measured sidereal rotation period of 12.76137 hours and the most likely pole orientation, $\lambda = 78^\circ$ and $\beta = +58^\circ$. Axial ratios are well constrained at $b/a = 1.239$ and $c/a = 0.819$.

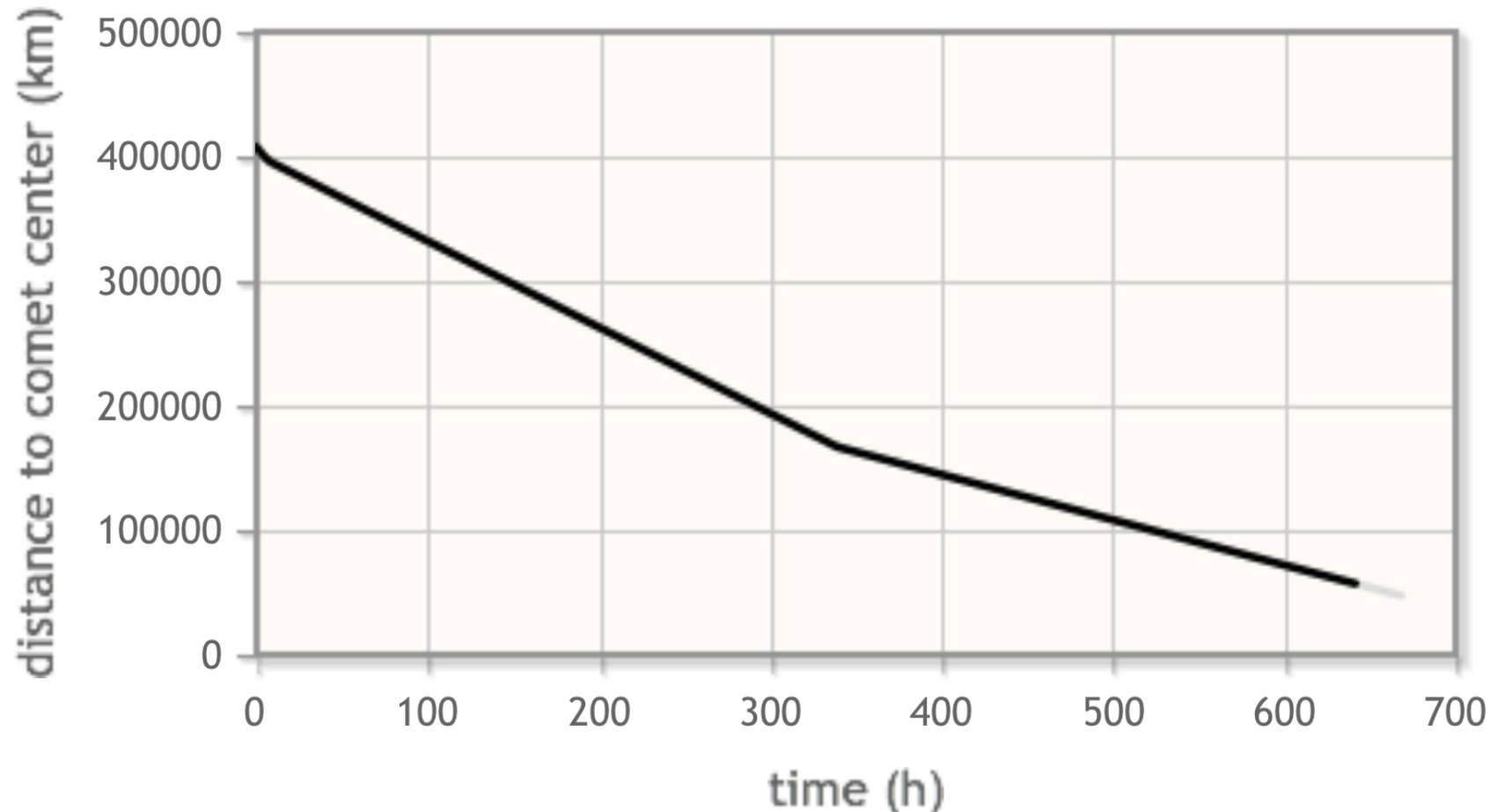


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afstand van de komet in juni 2014

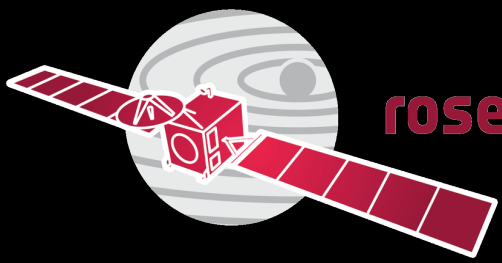
$\text{lat} = 47.9$





Navigation Camera NAVCAM1 20140608

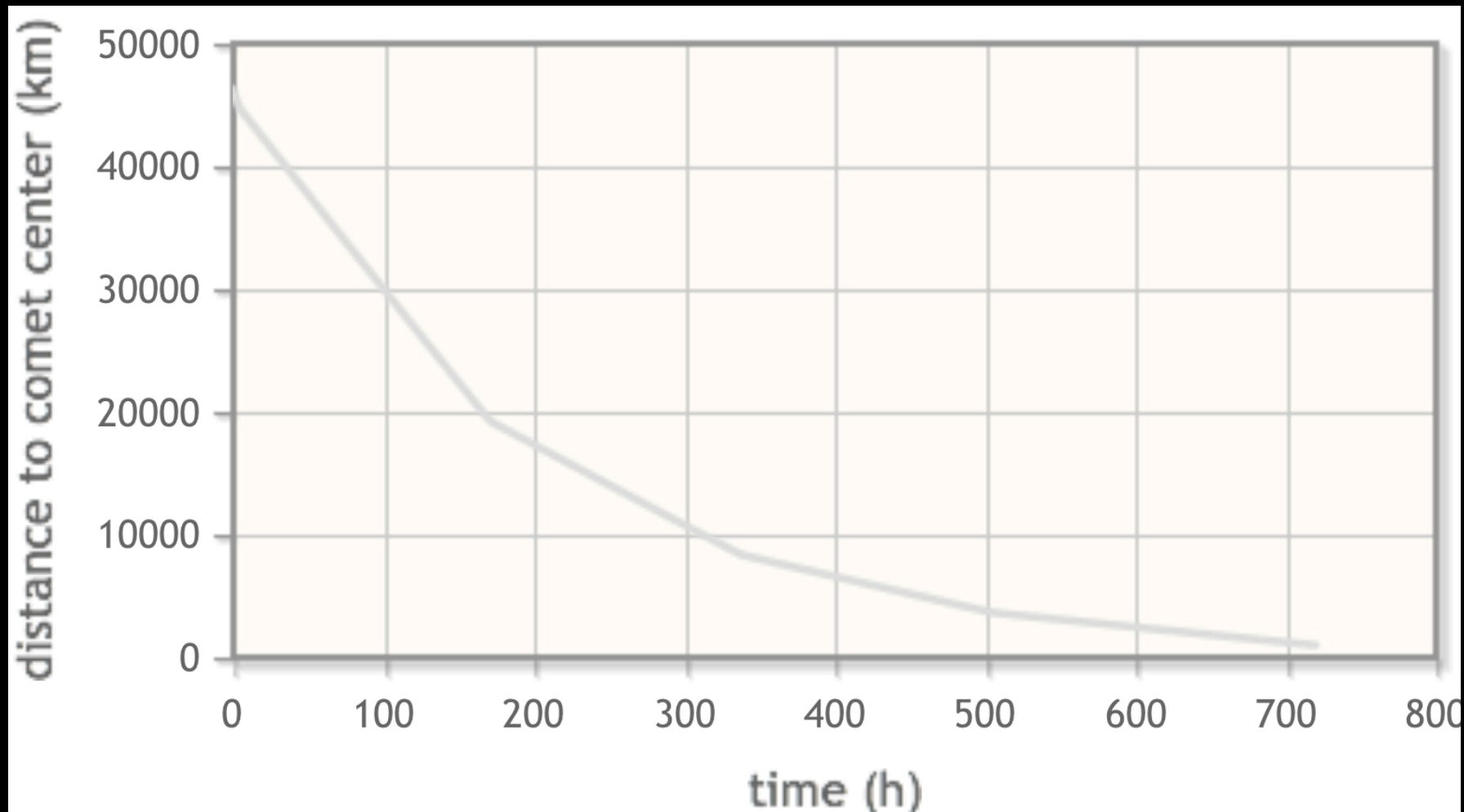




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afstand van de komet in juli 2014

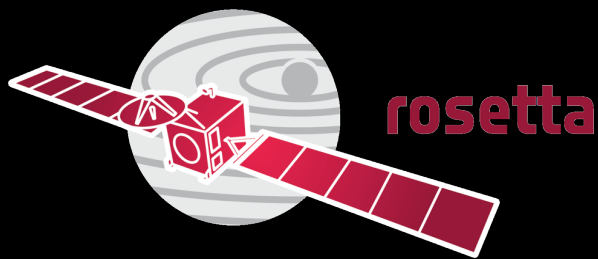




Navigation Camera NAVCAM1 20140703

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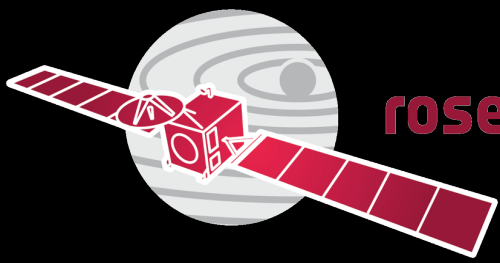




...

It was clear at the end of July 2014, that parameters and models needed adaptations and additional validation ...

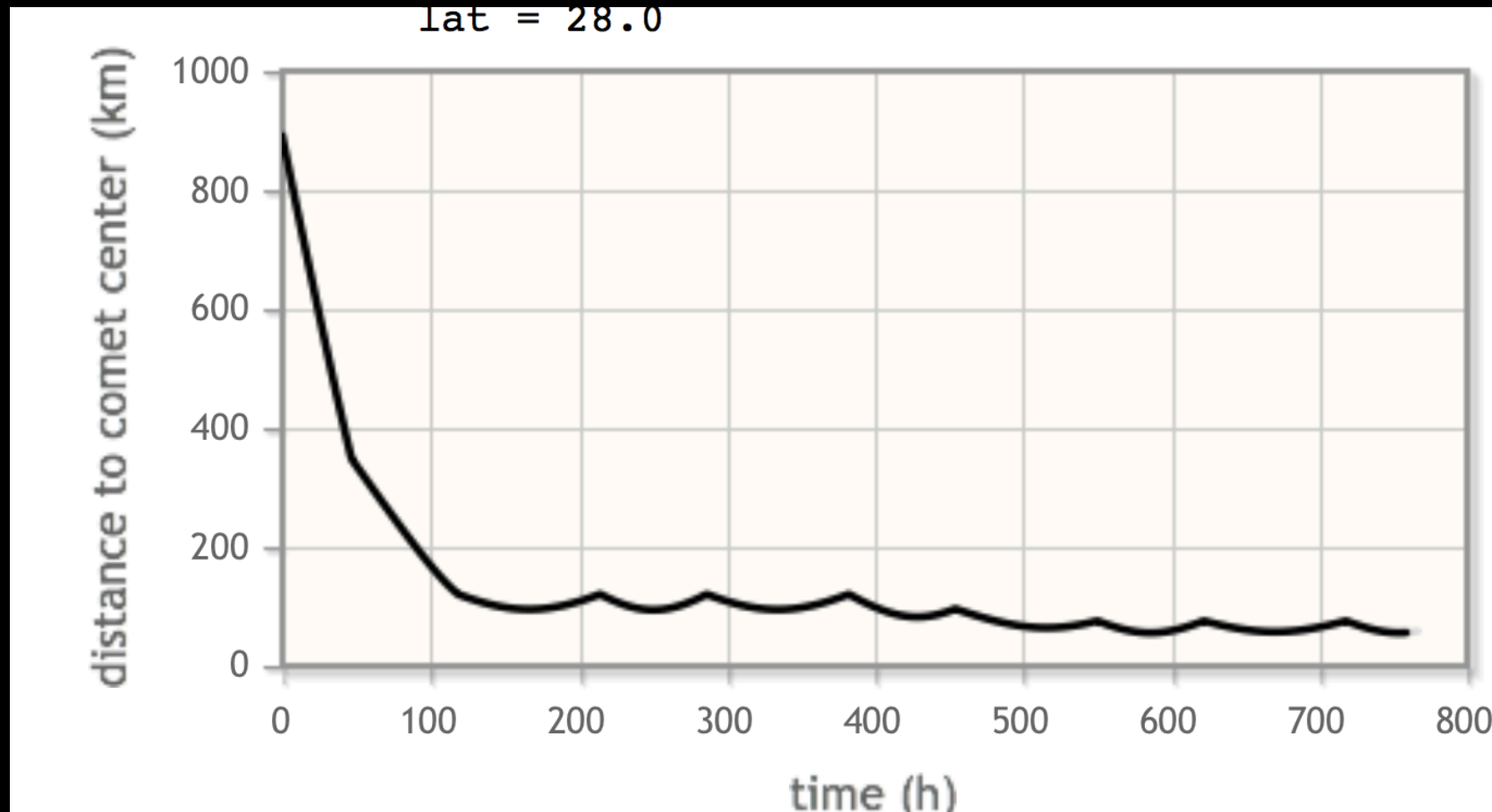
... but we did not yet have the full knowledge ... and decided to go on ...

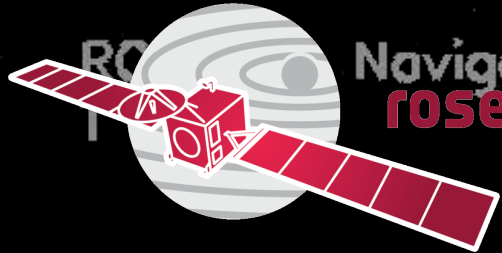


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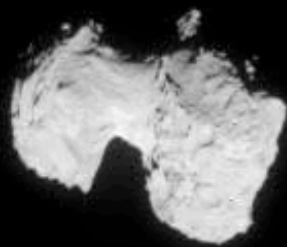
afstand van de komet in augustus 2014

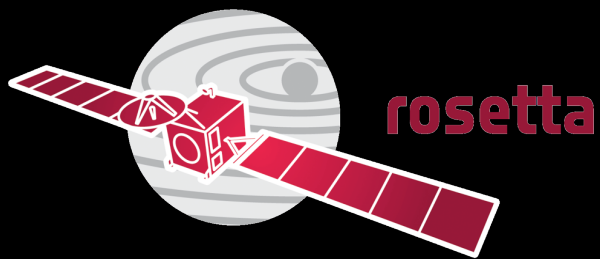




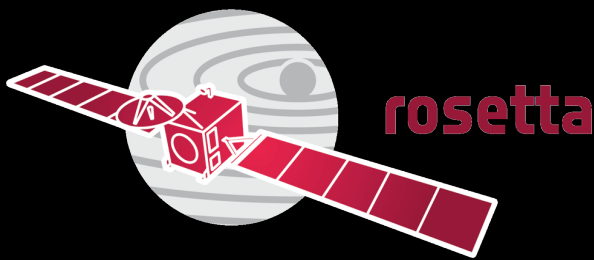
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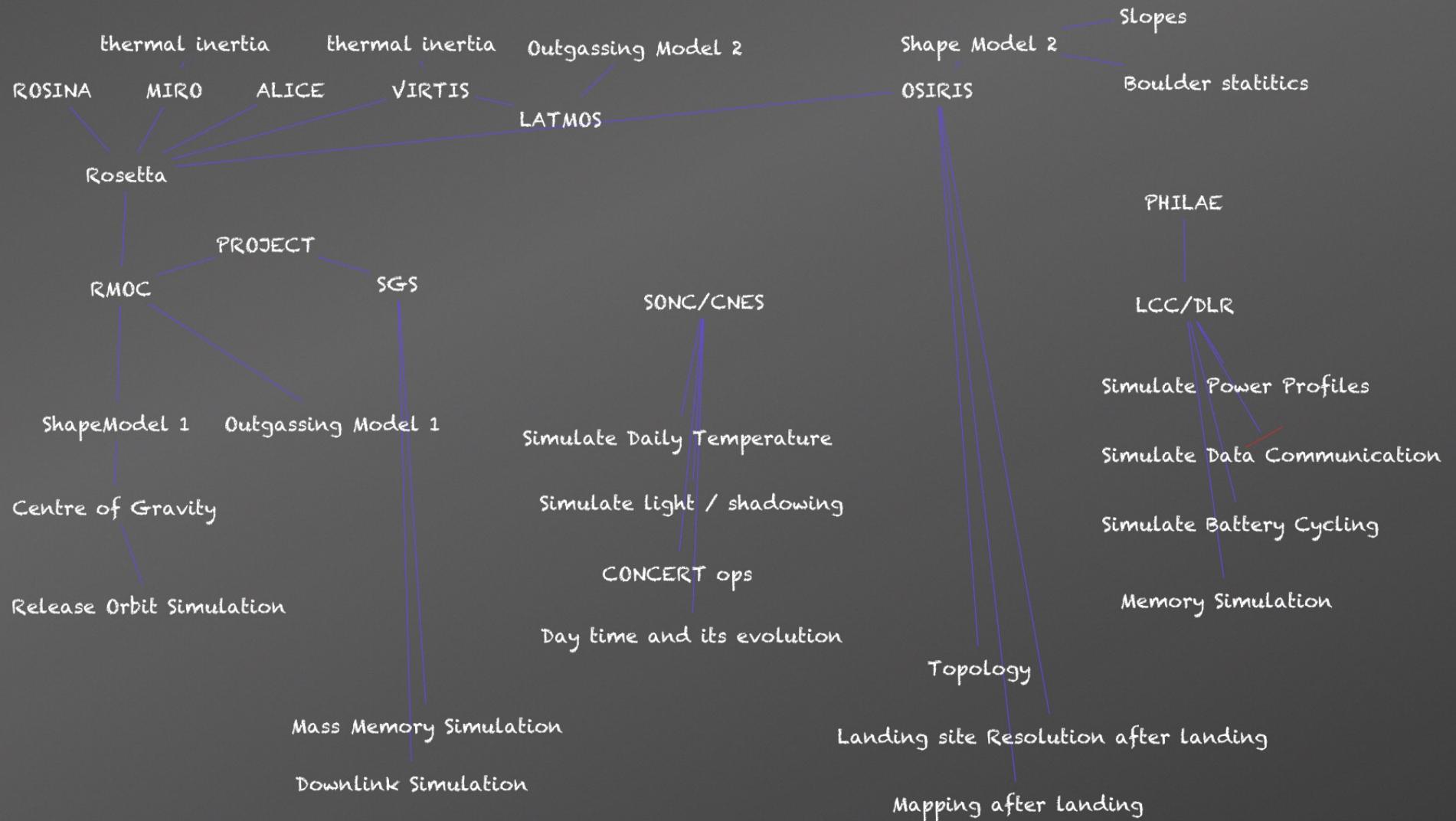
**... with limited information, we
went into the landing site selection meeting on 4-5 August
2014 ...**



status eind augustus:

- > Gaida detecteert de eerste stofdeeltje (1 augustus 2014)**
- > Pre-selectie van 10 landingsplaatsen (21 augustus 2014)**
- > komeet massa: oorspronkelijke raming: 3×10^{12} kg, nu 1×10^{15} kg**
- > Radius: oorspronkelijke raming: ~ 2.1 km, nu $[0.6, 2.8]$ km**
- > Dichtheid: 3 modellen werden gebruikt: $100/370/800$ kg / m^3**
- > Obliquity/schuinheid is constant: 53 graden \rightarrow CoG?**
- > CG buiten FOV van NavCam (22 augustus 2014)**
Selectie van 5 landingsplaatsen (24 augustus 2014)
- > Comet activiteit schatting en overeenstemming over wat er draait in LTP4 te vliegen (29 augustus 2014)**
- > Thermische kaart op komeet uit OSIRIS (29 augustus 2014)**

Name	Site	Long. / Lat.	Illumination 11/11/14	Illumination 01/11/15	Landing Scenarios	Dynamic Slope	Consent Color	Risk
LA_10	A	174 / 72	8.1	7.6	O1 – 6:30-8:3	15	Red	
In_Crater_2	B	353 / 5.2	6.4	6.3	O1 – 6:30-8:3	11	Green	
LA_16/HBO_4/ LA_12	C	205 / 26	7.6	7.2	O1 – 4:30-6:3	20	Yellow	
LA_14/HBO_3	D	134 / 45	10.3	9.1	O1 – 4:30-8:3	5	Yellow	Dispersion ellipse
LA_15/HBO_2	E	(78 / 24)	11.4	9.5	O1 – 6:30-8:3	10	Green	
LA_17/HBO_5	F	318 / 22	9.7	8.8	O1 – 4:30-6:3	22	Yellow?	
LA_13	G	105 / 50	10.0	8.8	O1 – 6:30-8:3	20	Red	
Polar_Day	H	133 / 59**	12.48	12.48	None	15	None	
HBO_1	I	14 / 23	9.12	8.21	O1 – 4:30-6:3	10	Green	Dispersion ellipse
SFD_1	J	335 / 25				25		



S/C security

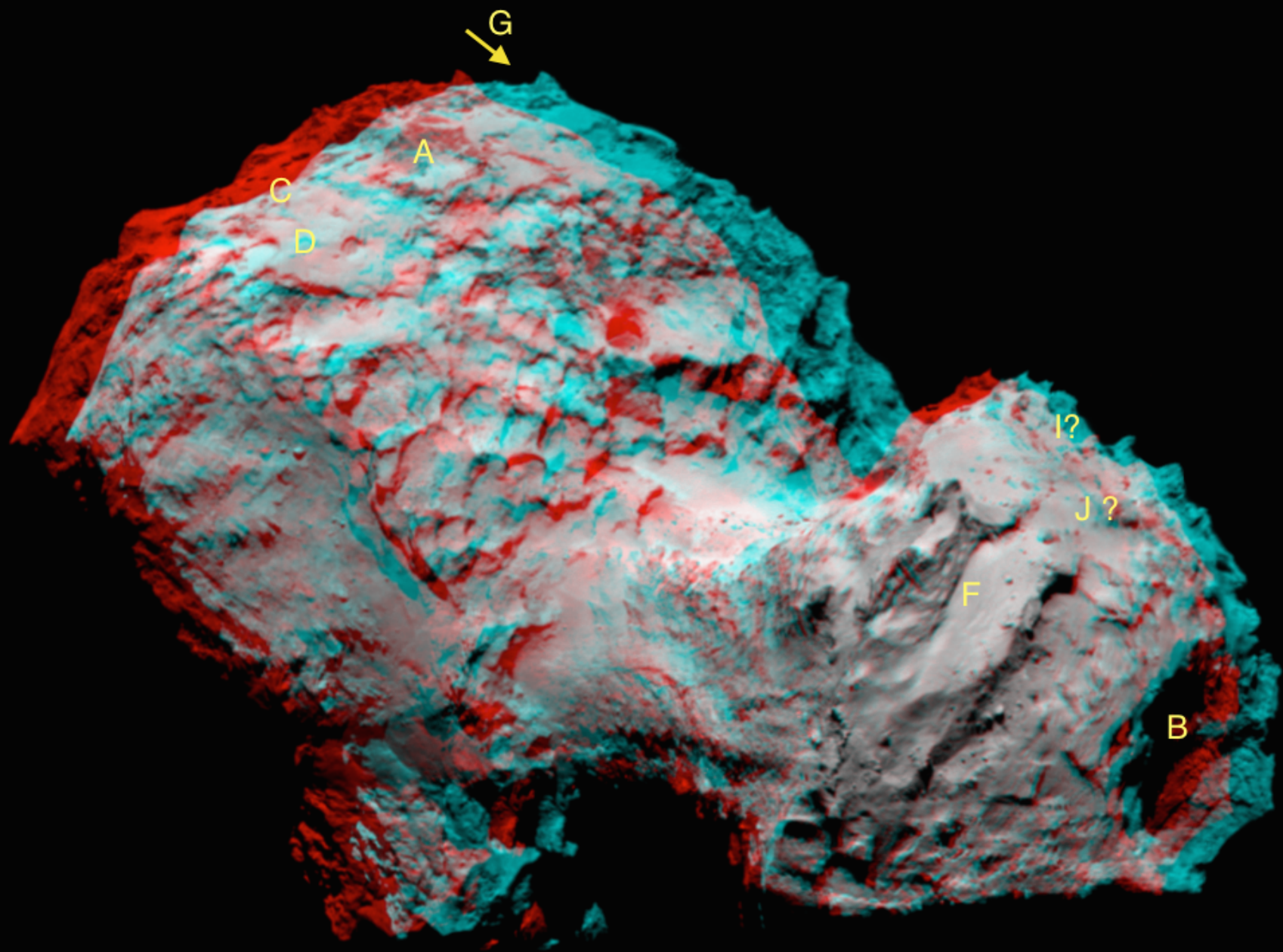
Landing Site (as a statistical parameter) Outreach

Science during Descent

First Science Sequence

Long Science Sequence

Lead Scientist

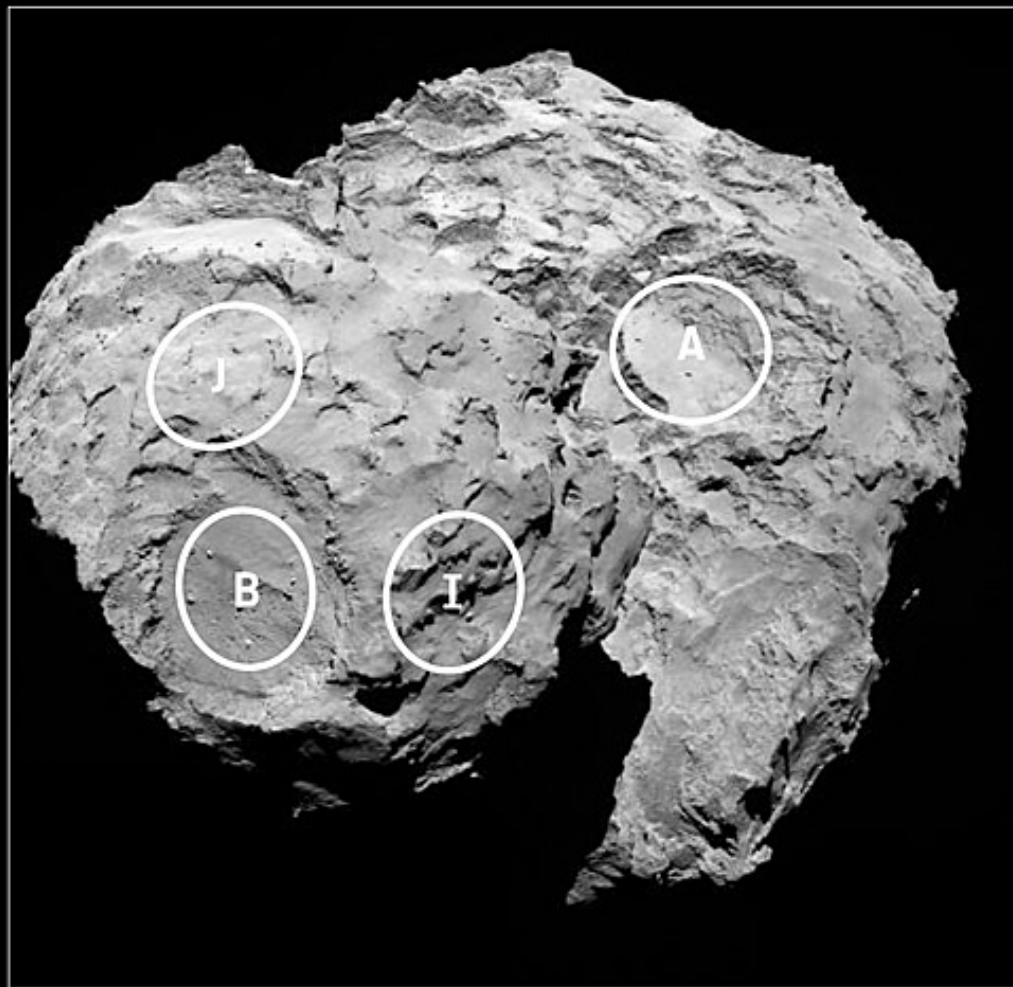


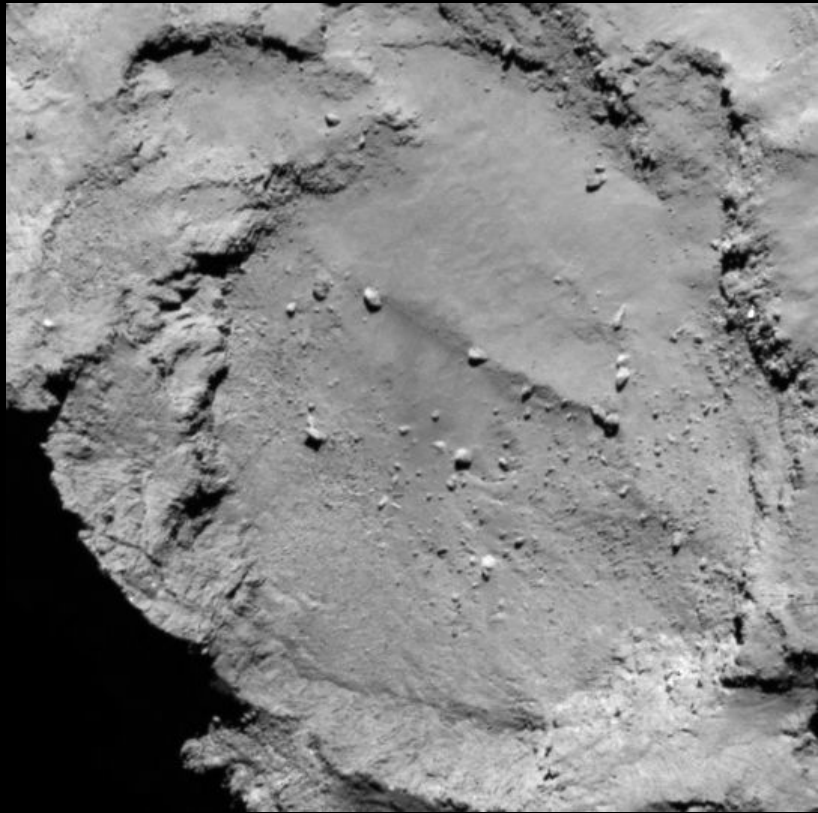
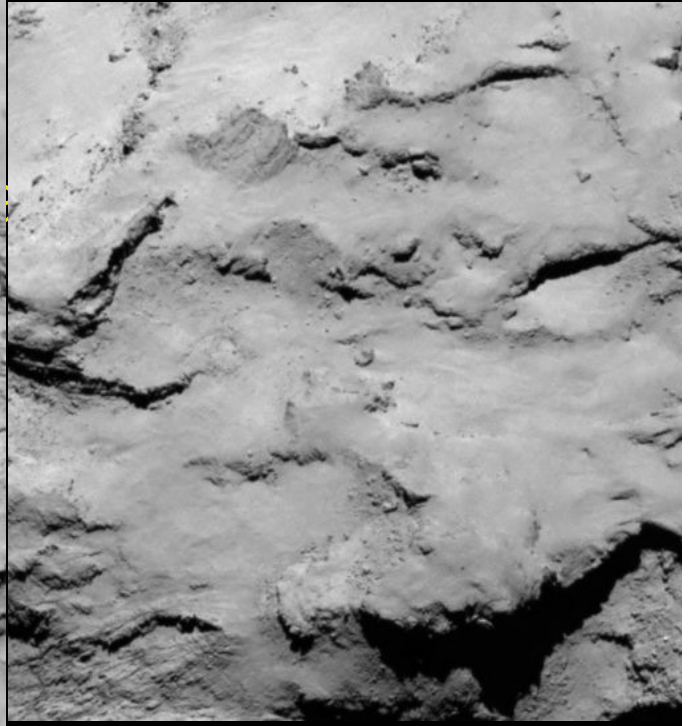
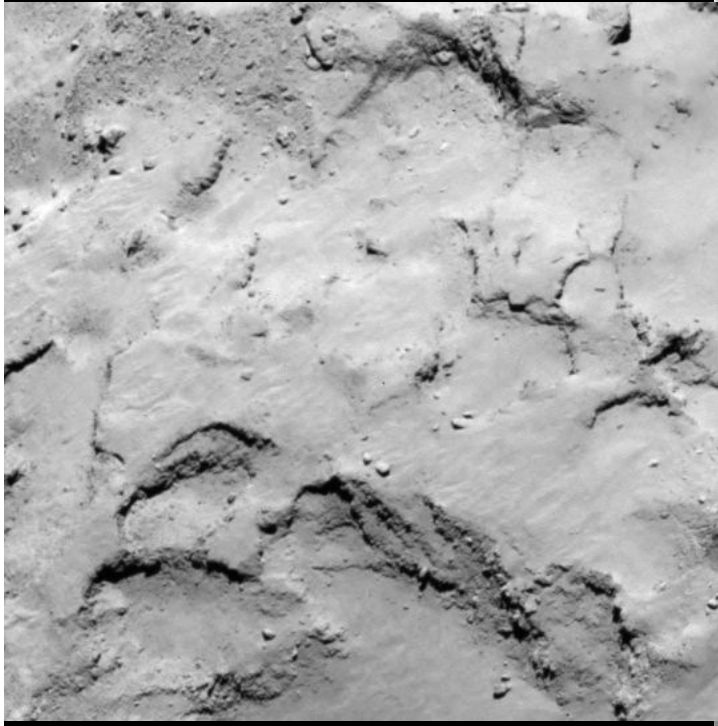


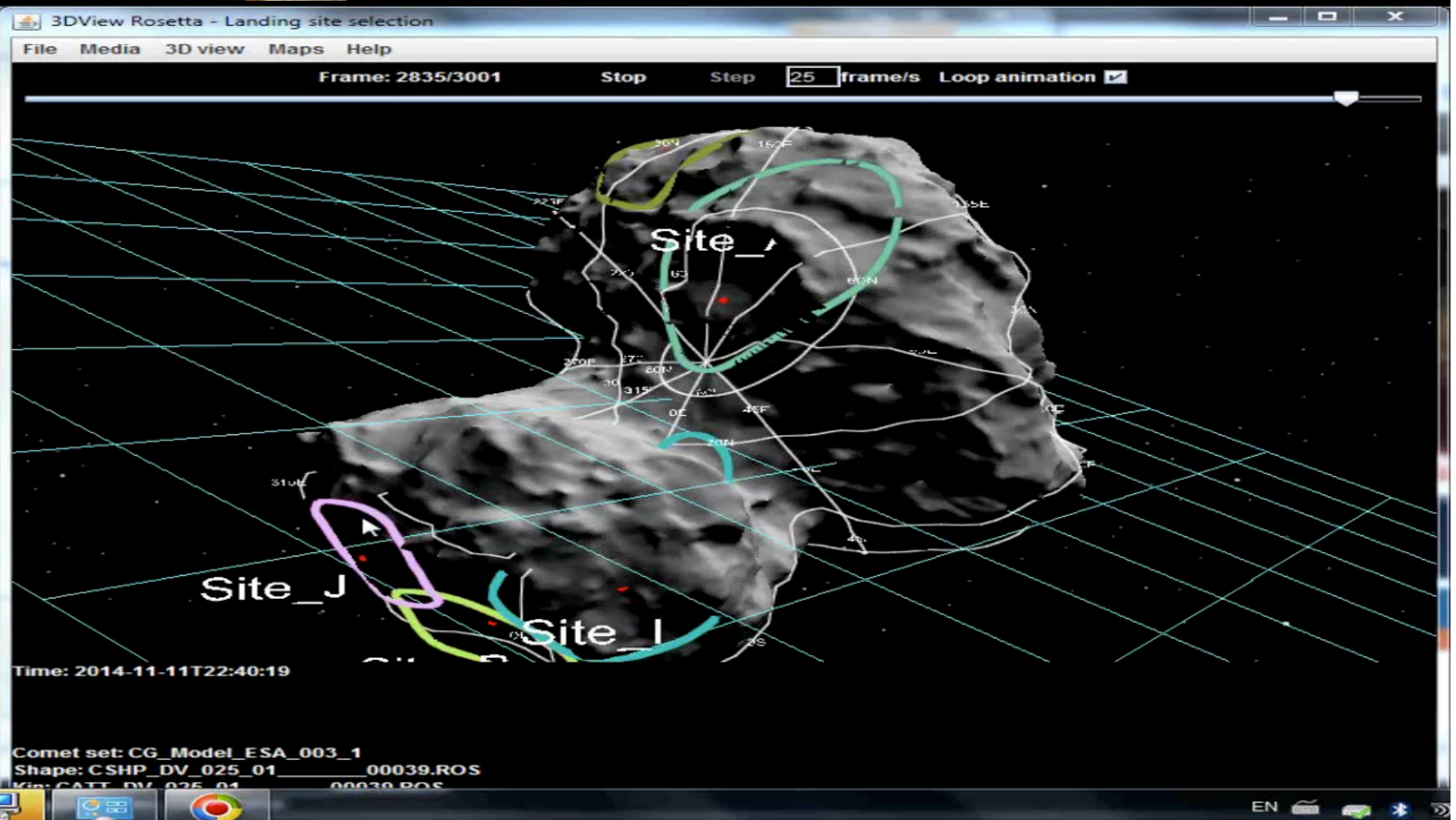
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the selected landing spots from 5 August 2014



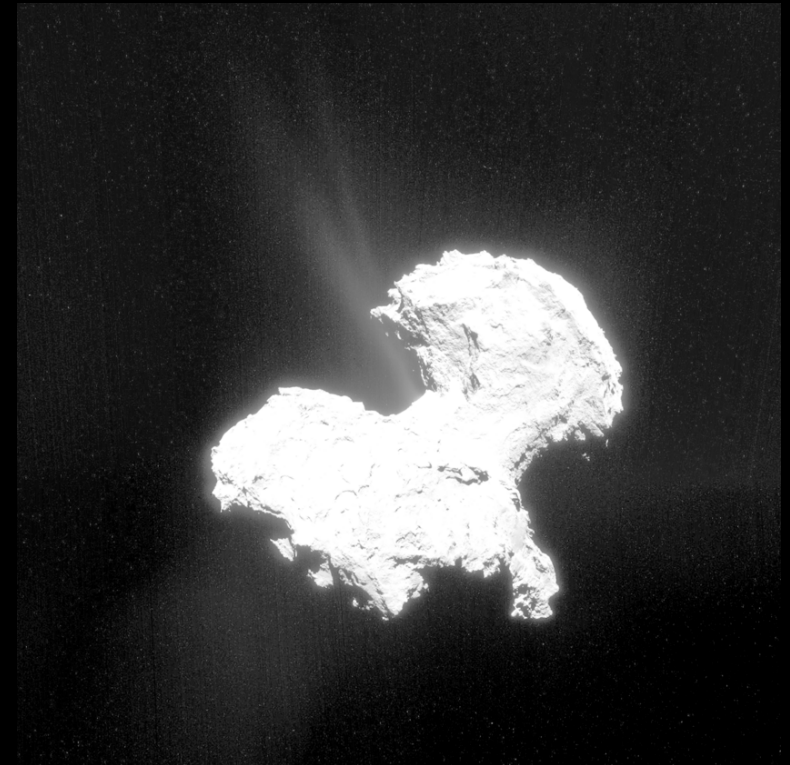
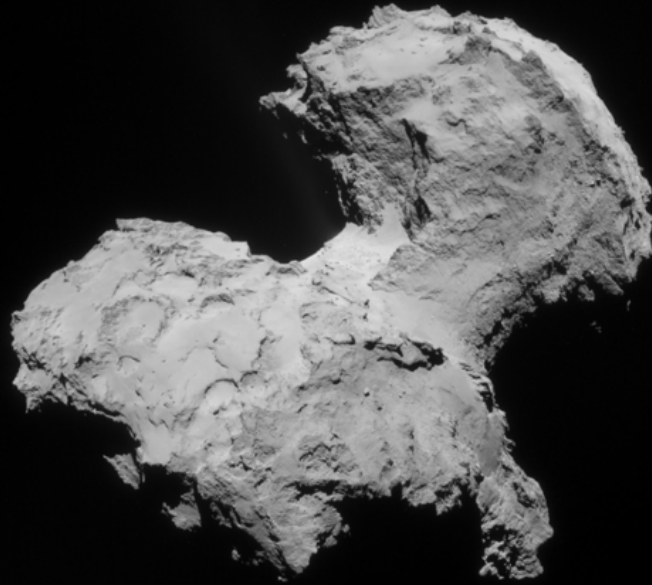


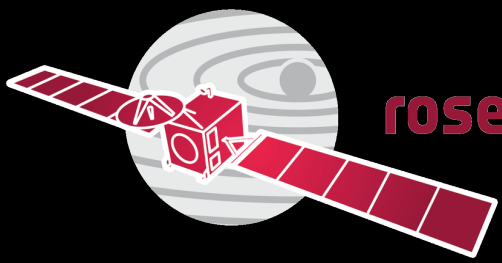




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3 september 2014

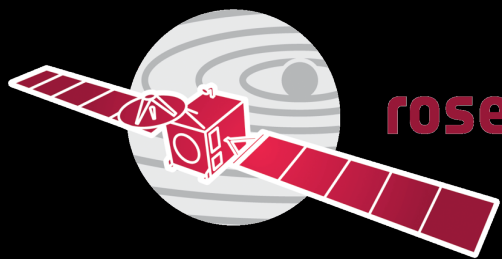




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from M. Malmer



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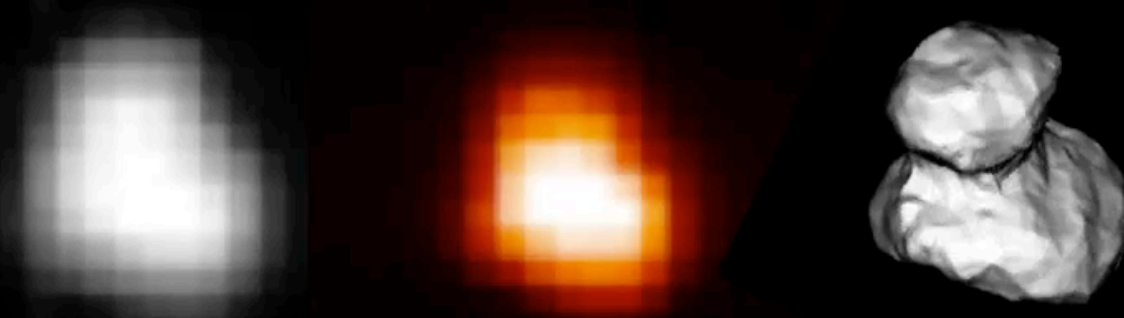


67P/CG NUCLEUS

2014-07-28T15:39

470 m/px

ESA/Rosetta/VIRTIS/INAF-IAPS/OBS DE PARIS-LESIA/DLR

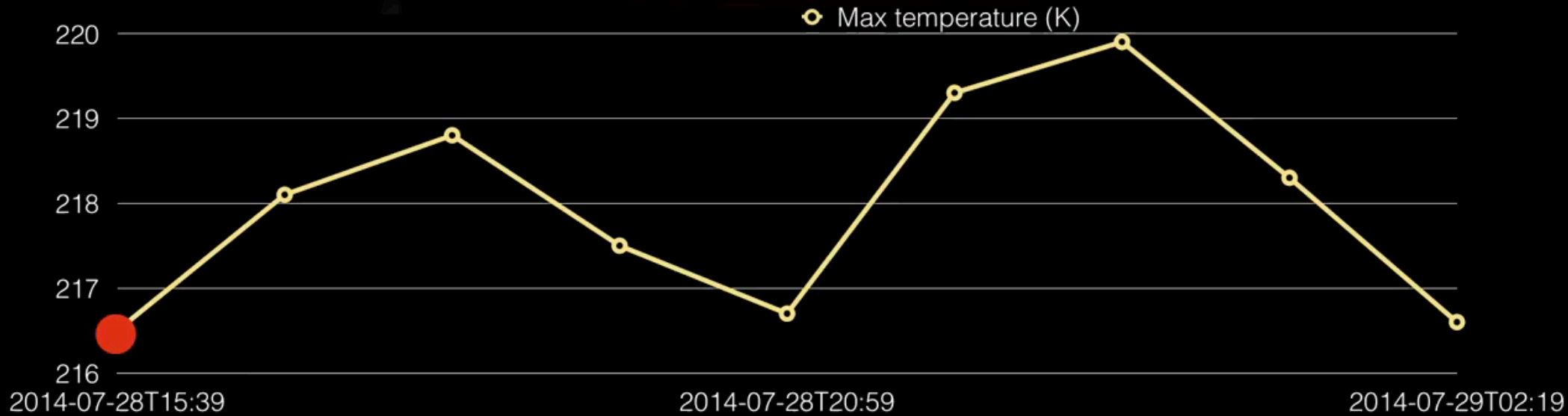


VIRTIS-M

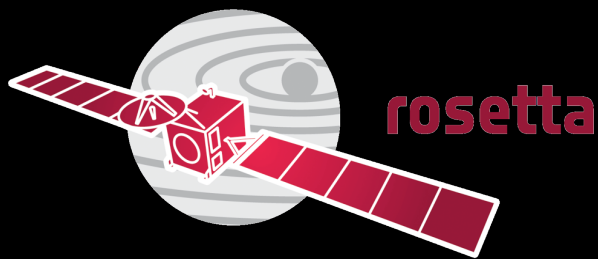
1.4 μm

5 μm

shape model rendering

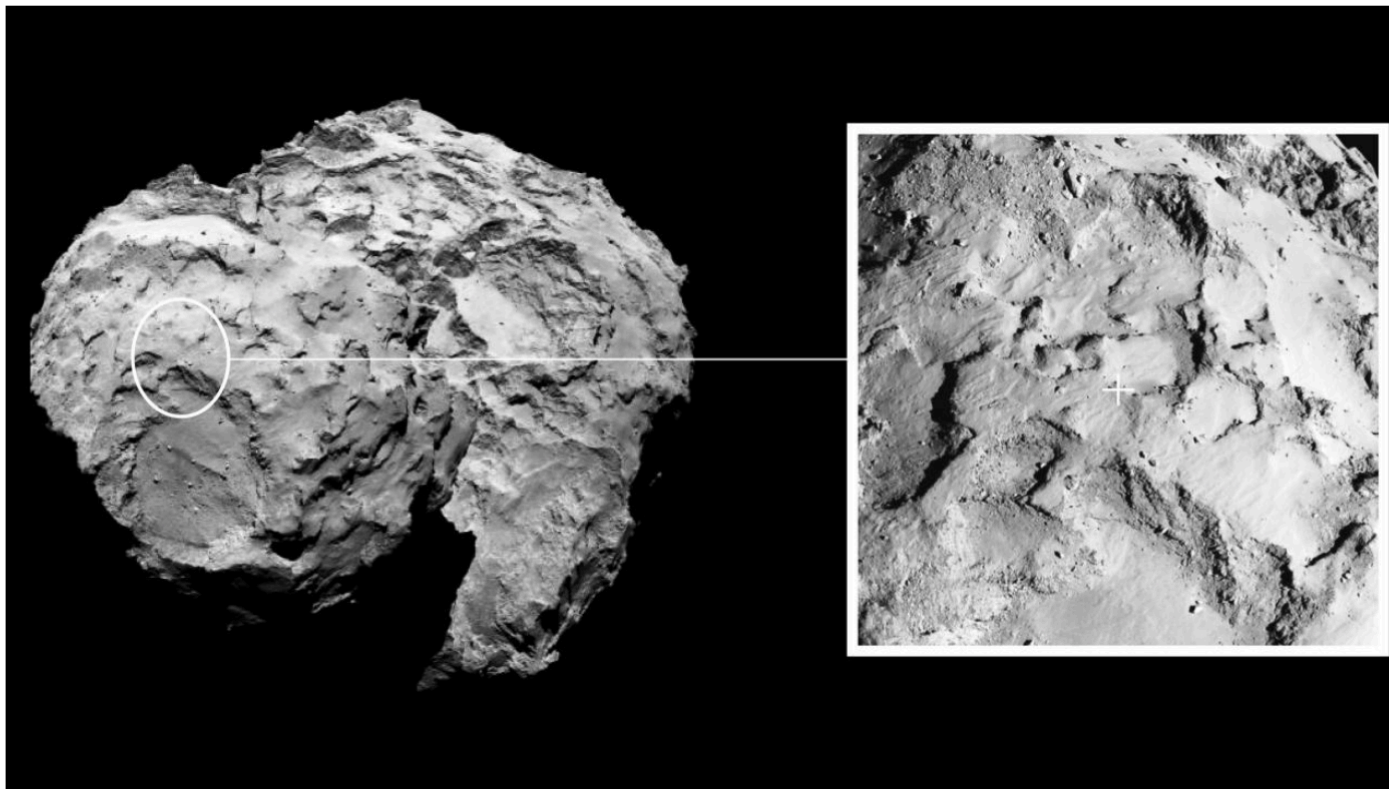






Selection of prime landing site on the week-end of 13/14
September 2014 (see ON_LSSP2 meeting
organization_Draft_270814)

SHARE

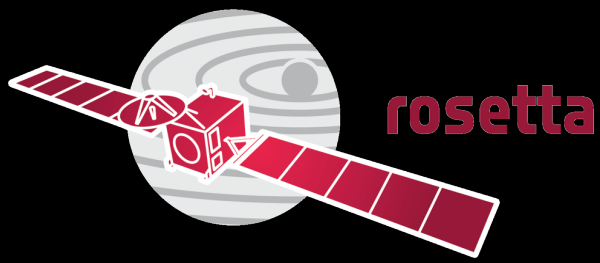


Circle marks landing spot: Lander hopes to avoid jets of gas and dust that could complicate descent.

ESA/Rosetta/MPS for OSIRIS Team MPS/UPD/LAM/IAA/SSO/INTA/UPM/DASP/IDA

Comet landing site picked: Rosetta to aim for the J spot

By **Eric Hand** | Sep. 15, 2014 , 8:15 AM



10 November 2017: GO/NOGO



esoc

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D-64293 Darmstadt
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MEMO

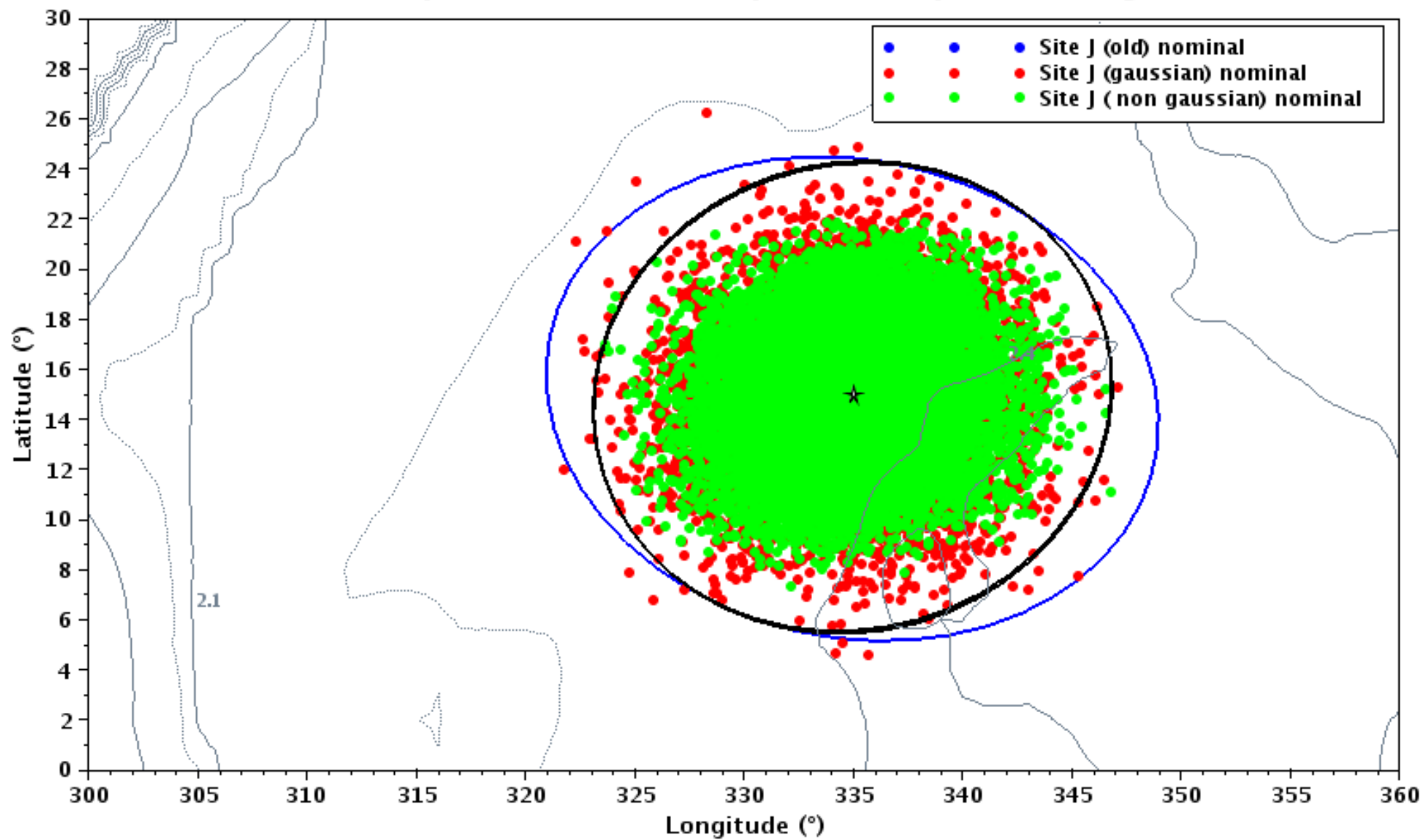
Date	09/10/2014	Ref	RO-ESC-ME-5508
From	C. Casas	Visa	
To	S. Ulamec, J. Biele, B. Paetz, M. Maibaum, K. Geurts, C. Fantinati, P. Gaudon, E. Jurado, C. Delmas, A. Blazquez (RLGS)	Copy	M. Ashman (RSGS); J. Fertig, V. Companys, R. Guilanya, R. Bauske, A. Accomazzo, S. Lodiote, A. Hubault, R. Kay (RMOC);

Subject: Assessment of prediction error for lander delivery

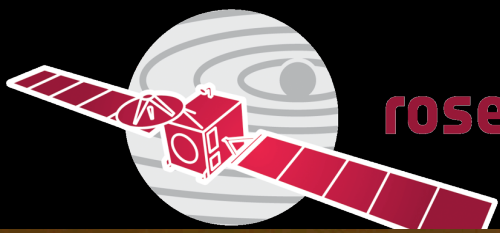
The navigation errors during the delivery sequence for both the orbiter and the lander for both site J (nominal) with strategy O1 and C (backup) with strategy O2 have been assessed as part of the LSSP-2 activity. The main results are specified in the LAFR delivered together with the LSSP-2 operational products. This memo shows results that have been requested by SONC but are not contained in the operational products.

The maximum expected navigation errors right before the pre-delivery manoeuvre for site J with strategy O1 (at a 30x30km orbit) are:

Comparison between dispersion ellipse for site J

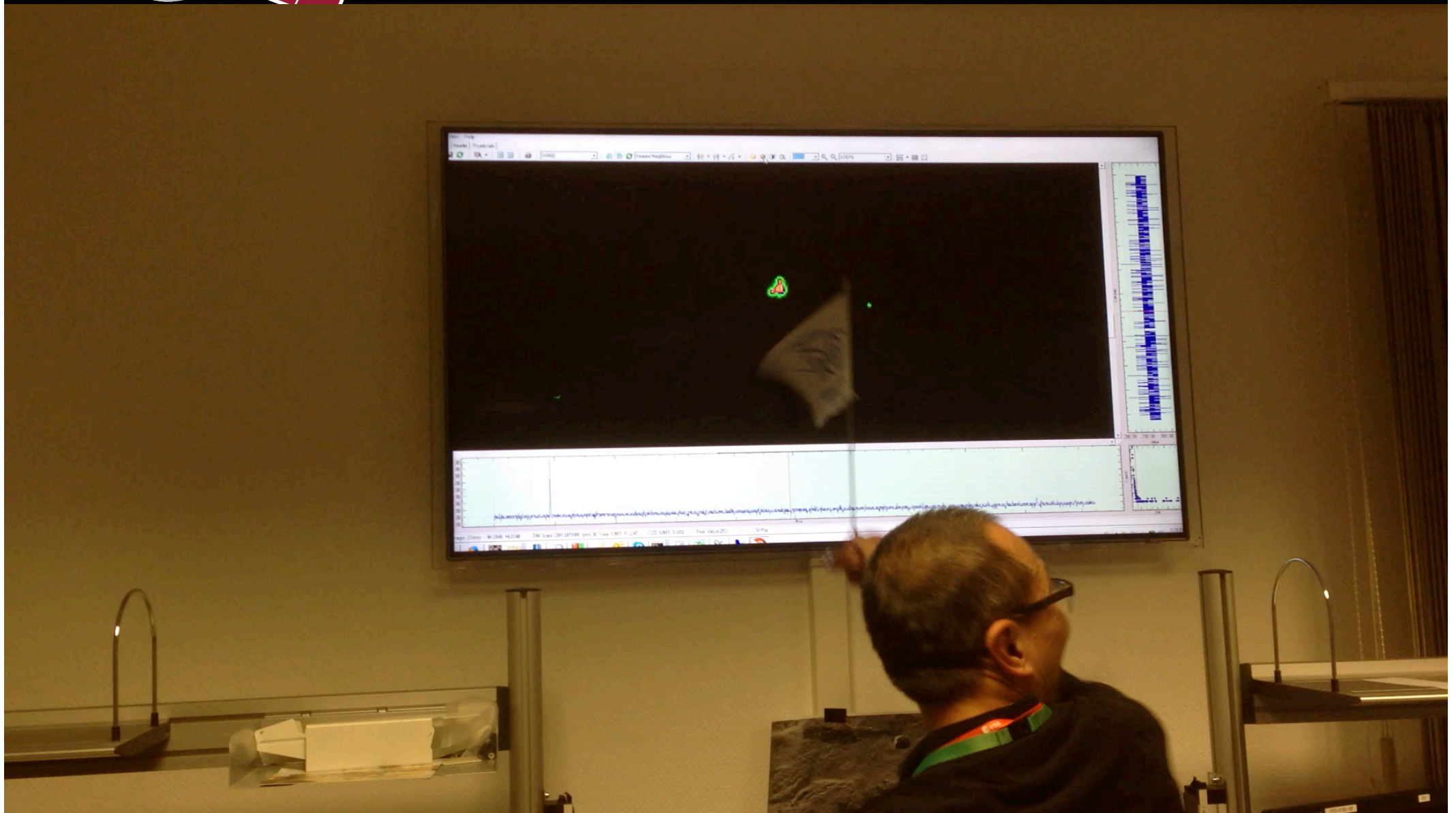


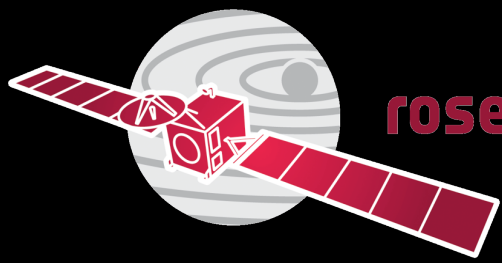
- show
LSSP-3_System_Report_L32_10102014



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Philae Landing



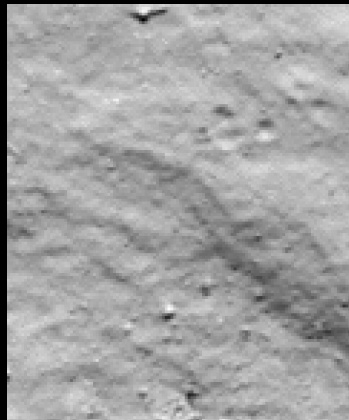
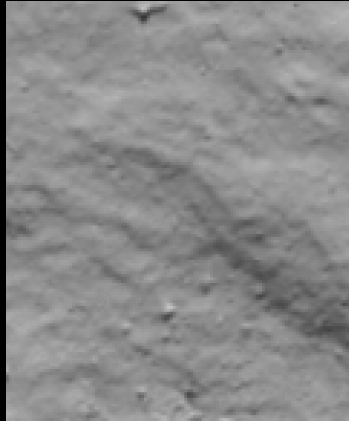


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Philae Landing



before



after



difference

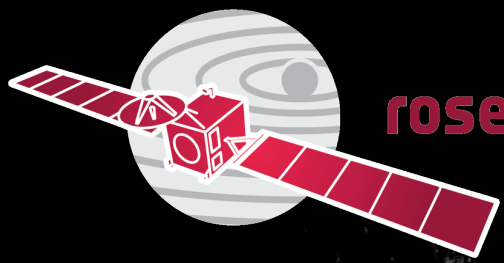


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count

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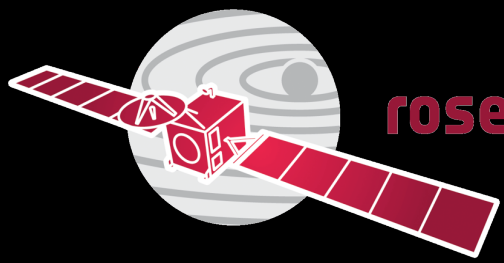
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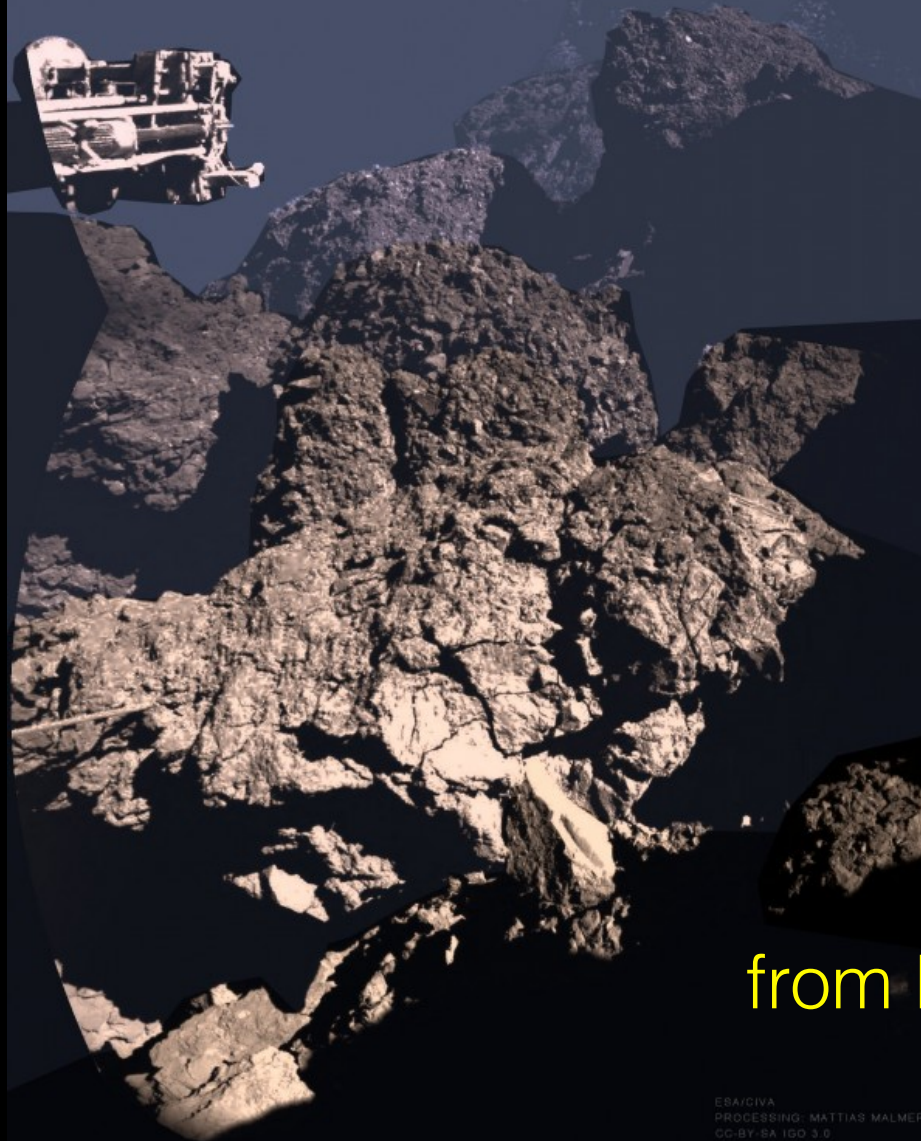
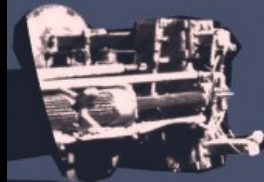


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from M. Malmer

European Space Agency

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